

Optical light curves of H-rich and H-poor Supernovae



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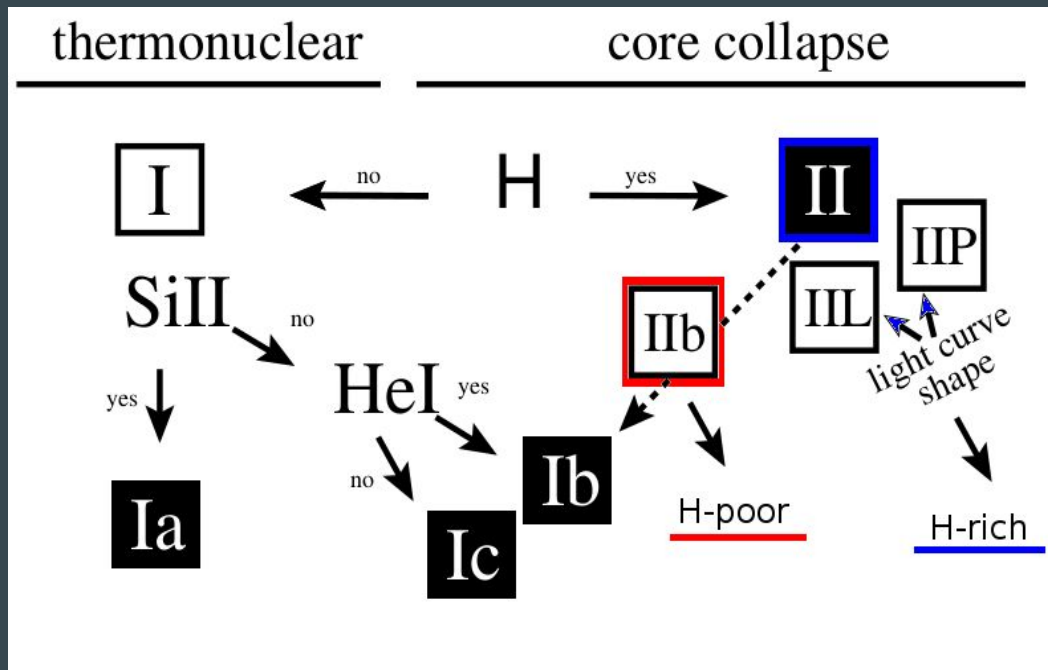
Director: Gastón Folatelli (IALP/FCAGLP)

Co-Director: Joseph Anderson (ESO)

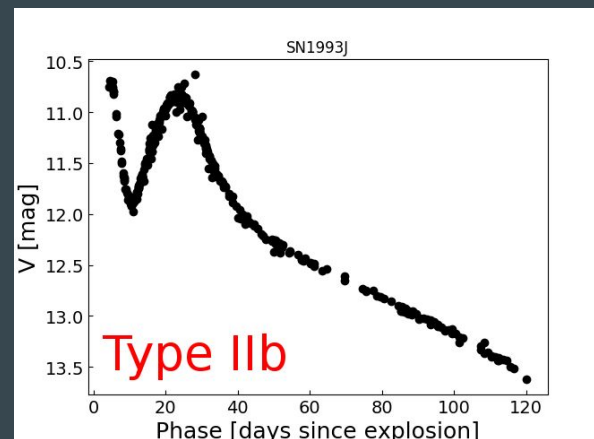
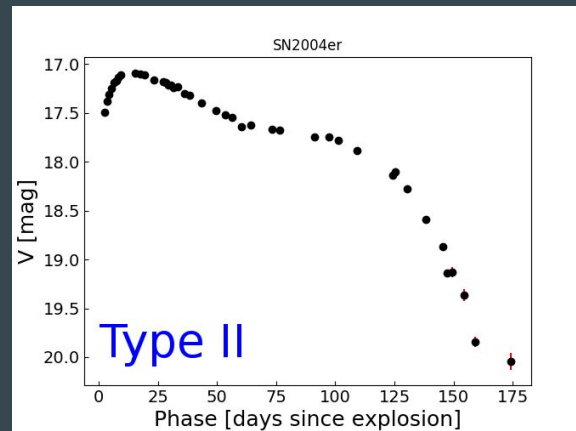
NidiaFest - Bariloche - 2018



SNe classification



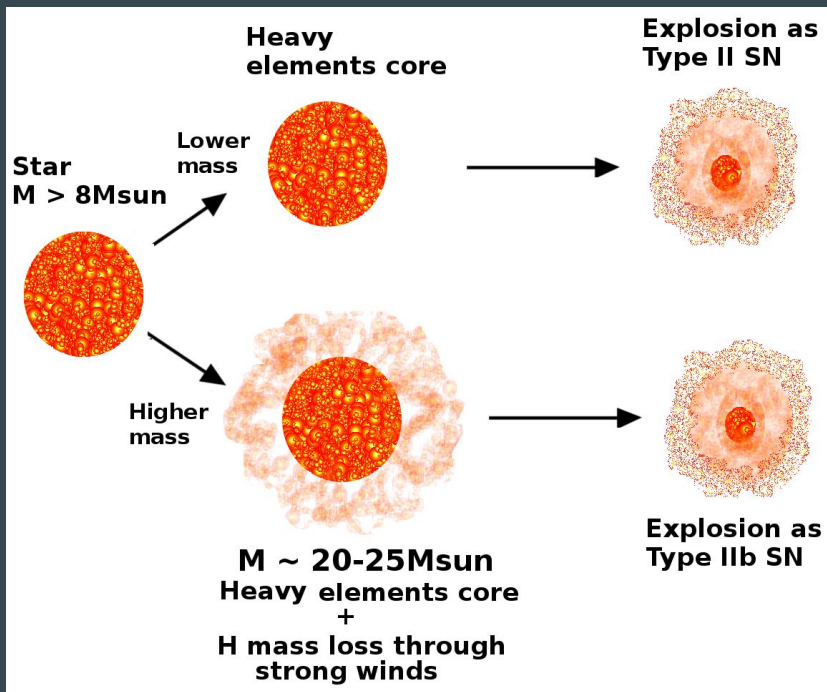
Turatto03



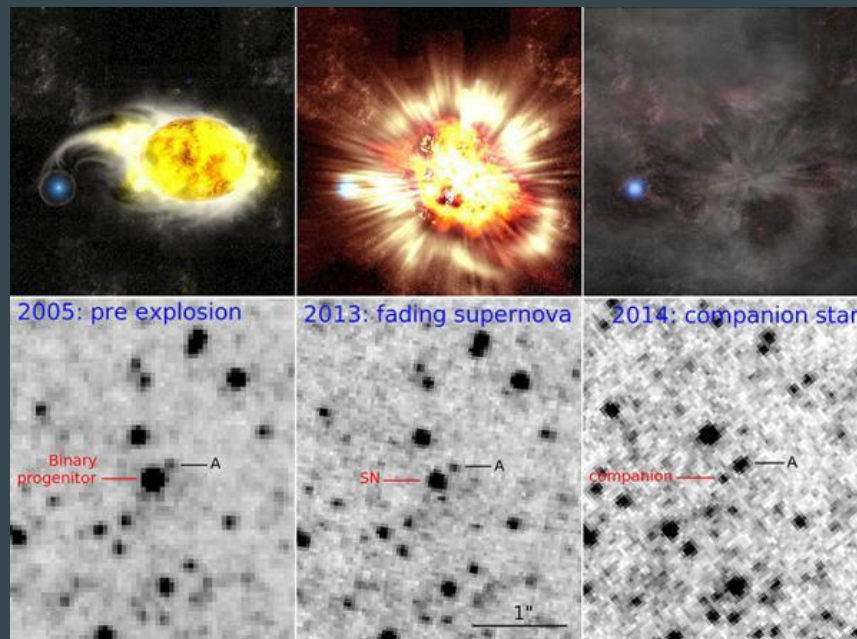
The core collapse progenitors question

Understanding progenitors of different SN types we can learn about how massive stars lose mass

Single progenitor scenario



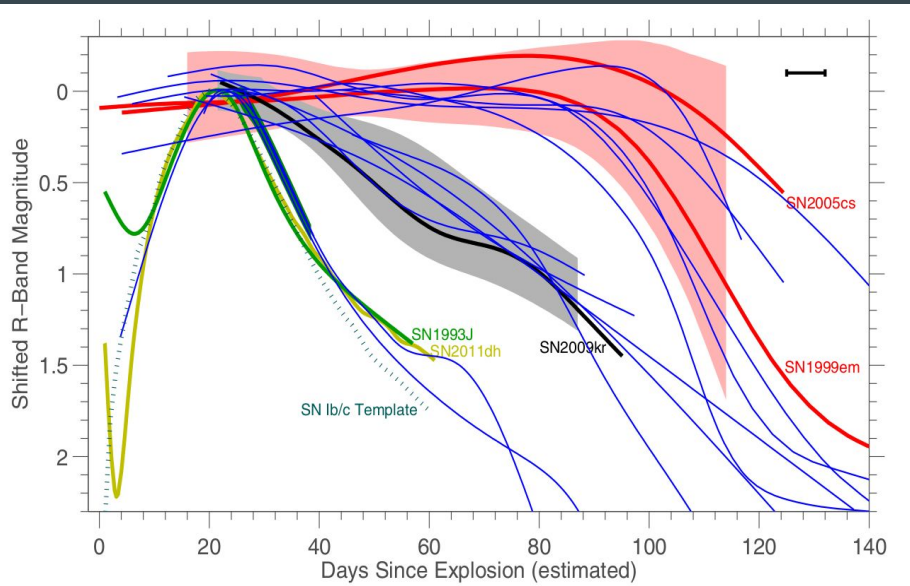
Progenitor in binary system scenario



Folatelli+14

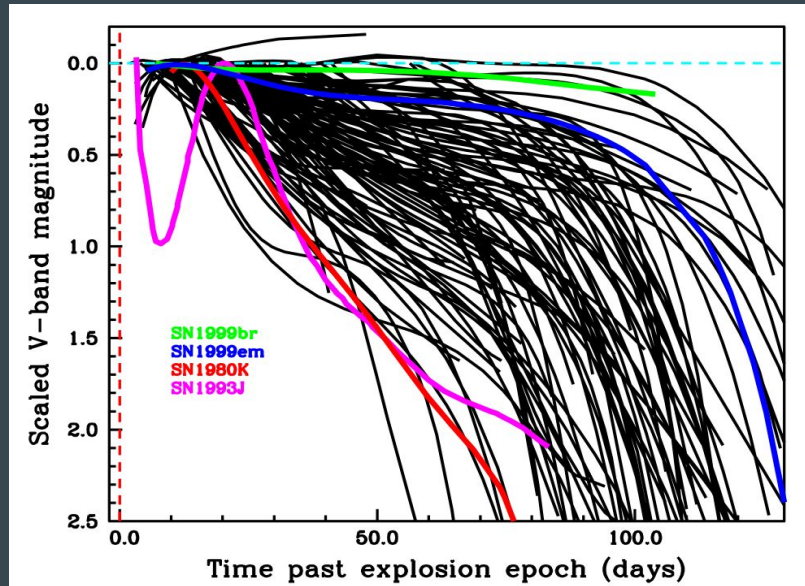
Previous studies

Arcavi+12



“... Type IIP, IIb, and IIL SNe might not be members of one continuous class...” (Sample: 21 SNe in R band)

Anderson+14



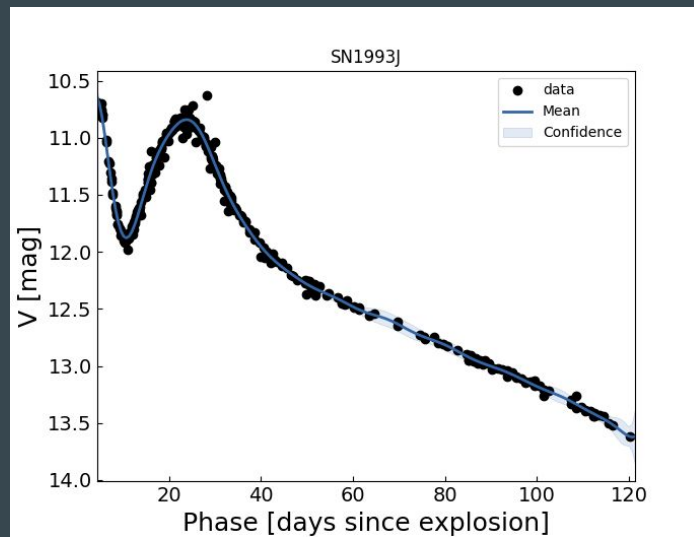
“A continuum of SN II V-band properties is observed in all measured parameters...” (Sample: 116 Type II SNe in V band).

Our work

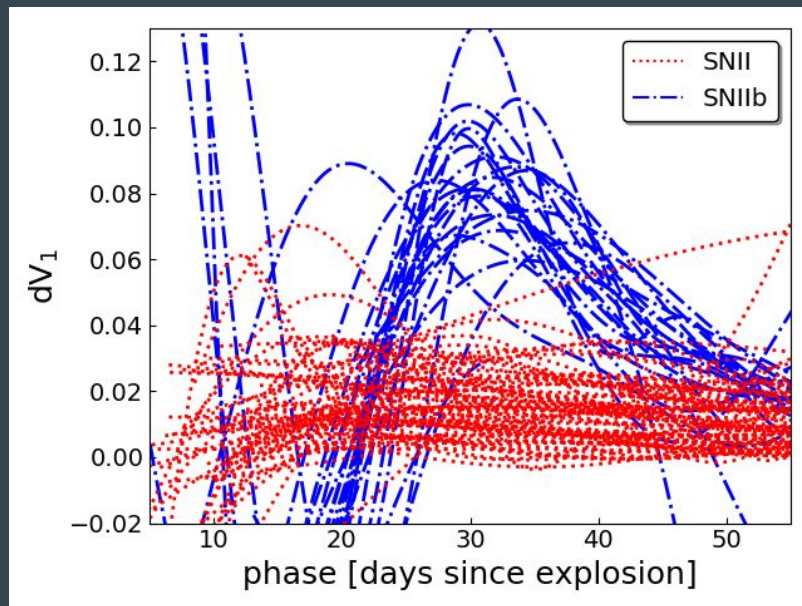
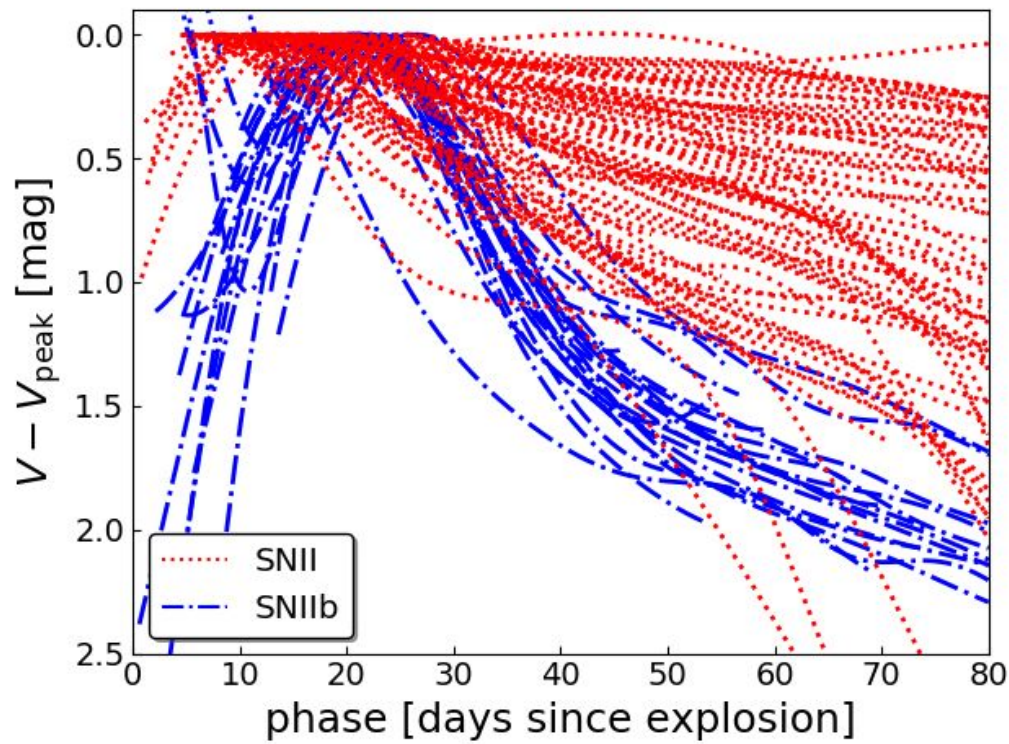
- **Goal:** Test the hypothesis of a continuum of physical properties between SNI_I and SNI_{Ib} by examining their observed properties.
- **Strategy:** Resample the light curve of Type II and Type IIb SNe in order to be able to define parameters that describe the LC morphology to compare both families.

→ **Data:** 73 SNI_I and 23 SNI_{Ib}.
B, V and r bands.
CSP+CATS+Others.

→ **LC Resampling:** Gaussian process via implementation of the Python library GPy.



LC morphology

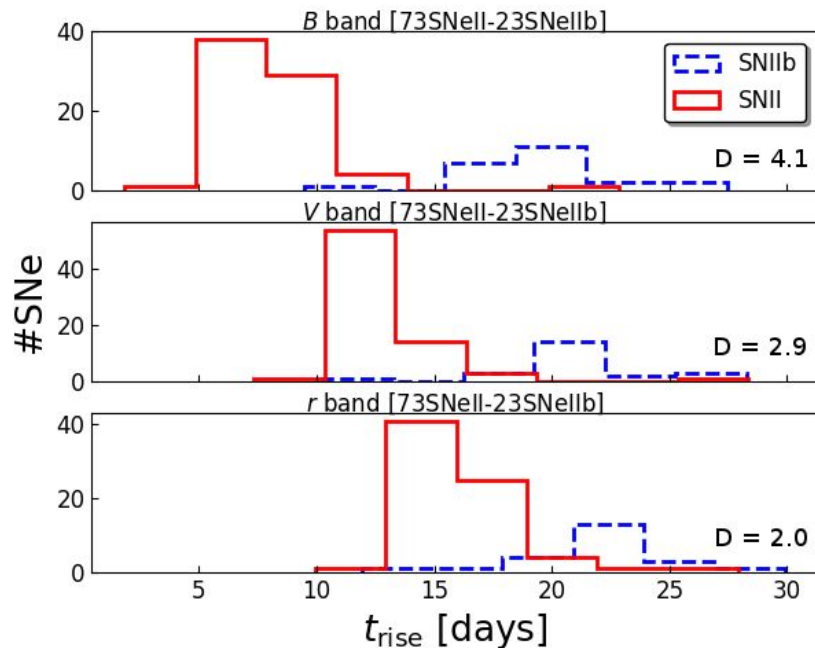


Parameters and results

- **Rise Time:** elapsed time between explosion and maximum light.

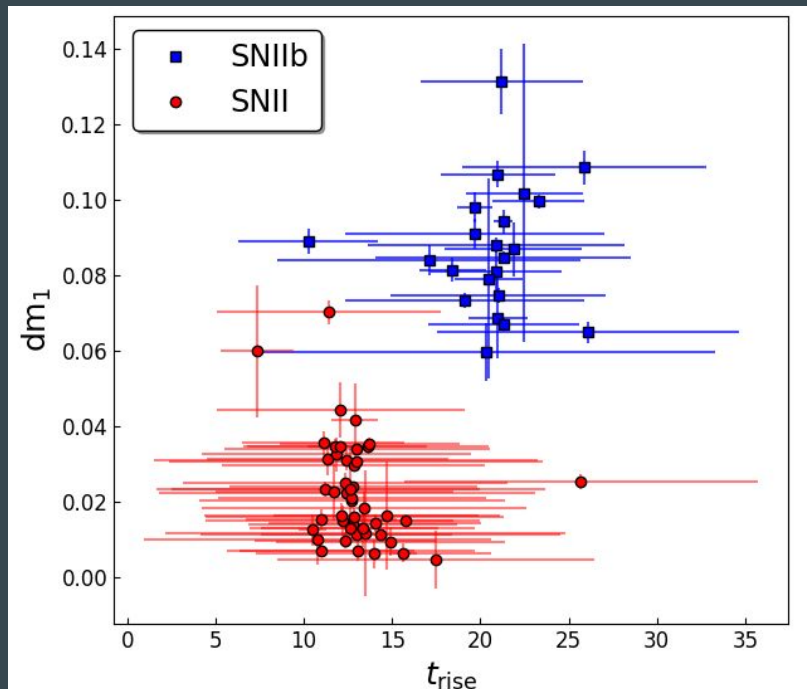
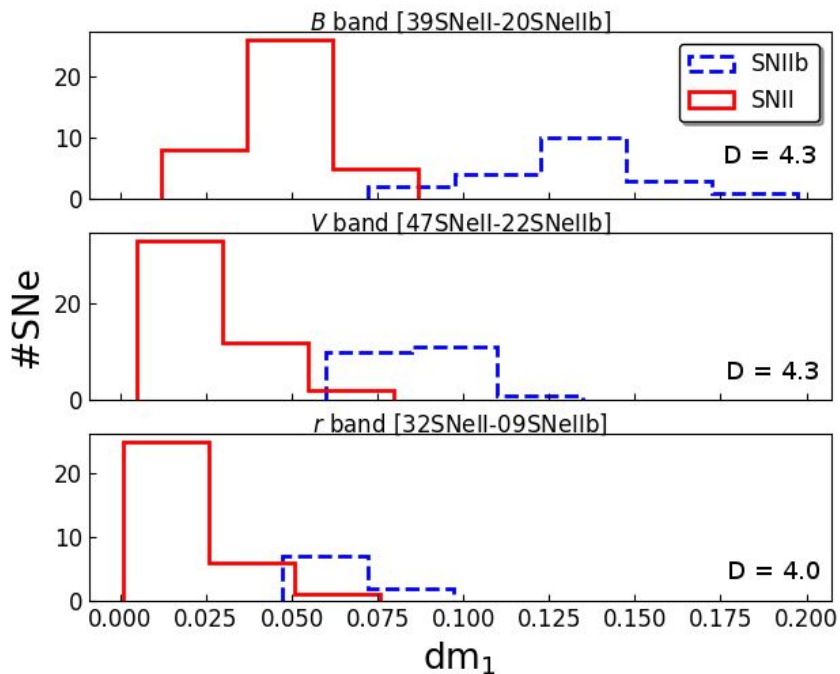
Weighted average of t_{rise} for every band.

Type	Band	Weighted average
IIb	<i>B</i>	19.0 ± 1.8
	<i>V</i>	20.9 ± 1.6
	<i>r</i>	21.5 ± 2.1
II	<i>B</i>	8.3 ± 2.0
	<i>V</i>	12.8 ± 2.4
	<i>r</i>	16.0 ± 3.6



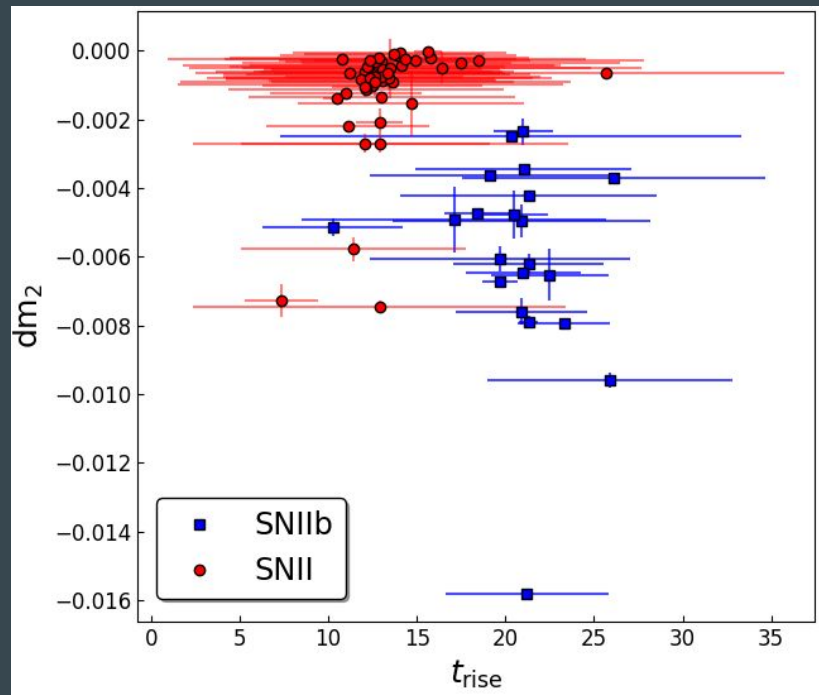
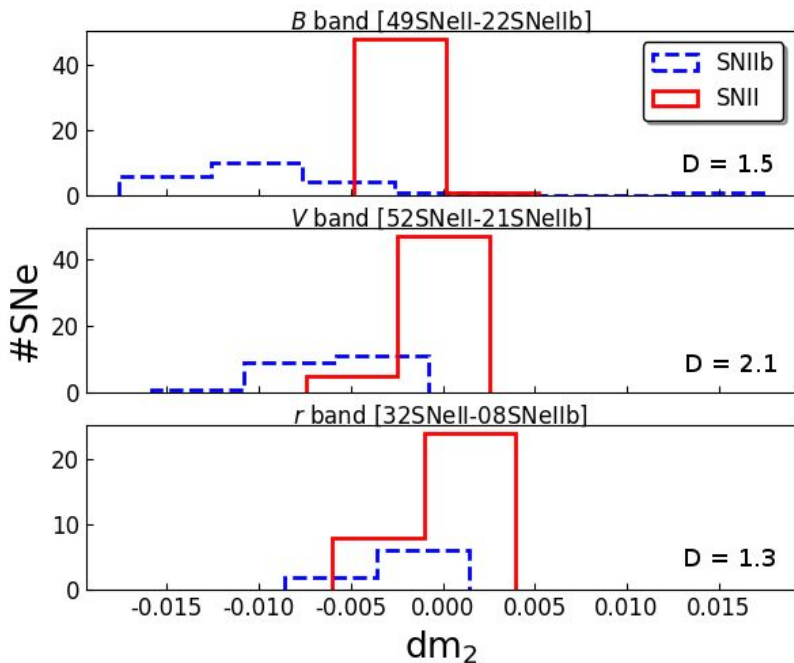
Parameters and results

- **First maximum of the first derivative after peak:** accounts for the steepness of the LC.



Parameters and results

- **First minimum of the second derivative after peak:** gives an idea of how much the curvature of the LC changes in the region where it changes the most.



Conclusions

- Overall we see both families well differentiated (in B_Vr), suggesting that there is no continuum of observed properties among SNI_I and SNI_{Ib}.
- We infer from the above that most likely the progenitor population is also not a continuum and both H-rich and H-poor SNe II might come from different progenitors.
- Our findings agree with the recent claims that SNe II form a single continuous family in terms of LC decline rate between what has been called II-P and II-L SNe.

