

High Contrast Imaging of Massive Stars

Alan Rainot
Hugues Sana & the CHIPS team
K.U. Leuven

5th November 2018
San Carlos de Bariloche, Argentina



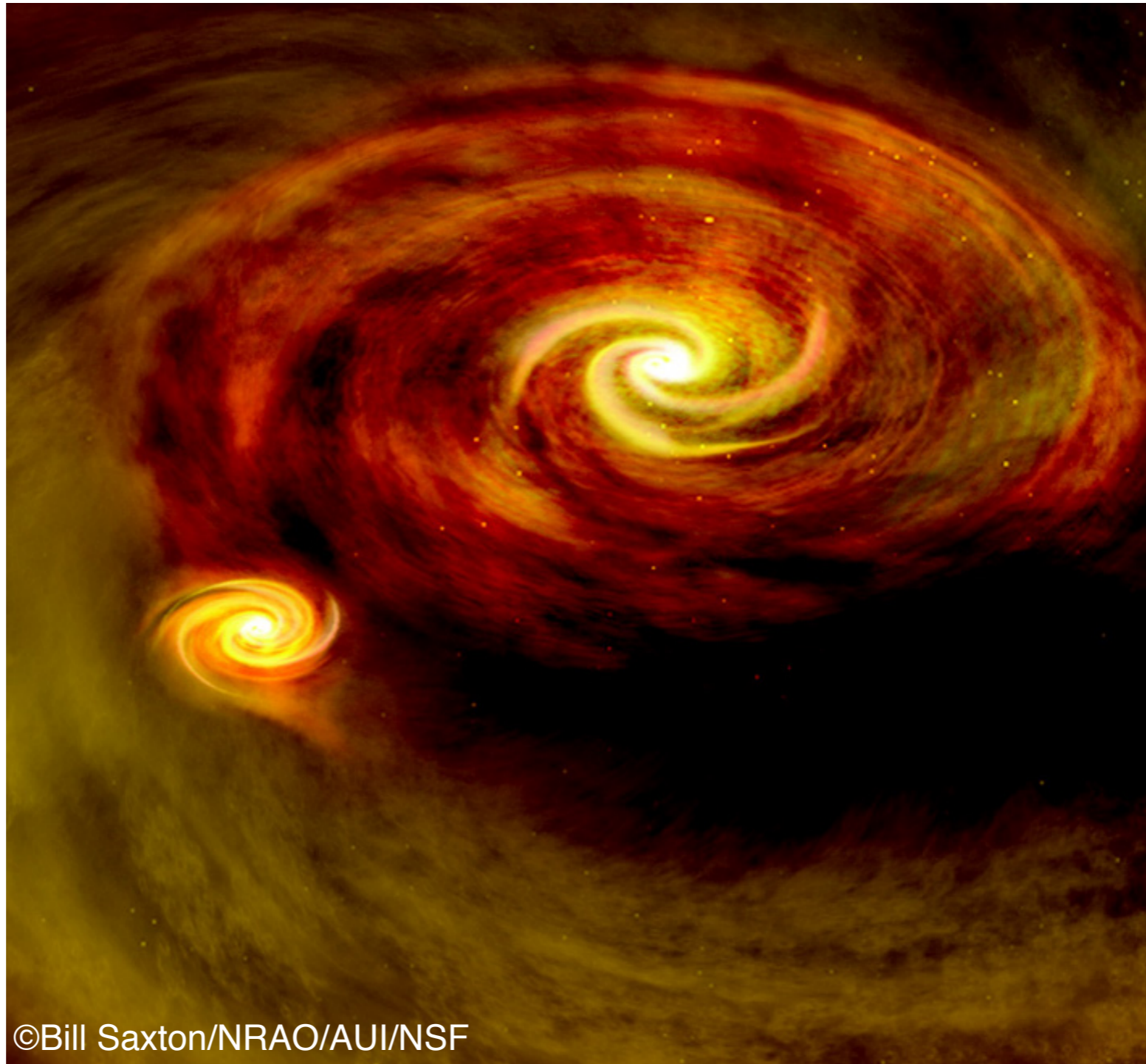
<https://arainot.github.io>

alan.rainot@kuleuven.be

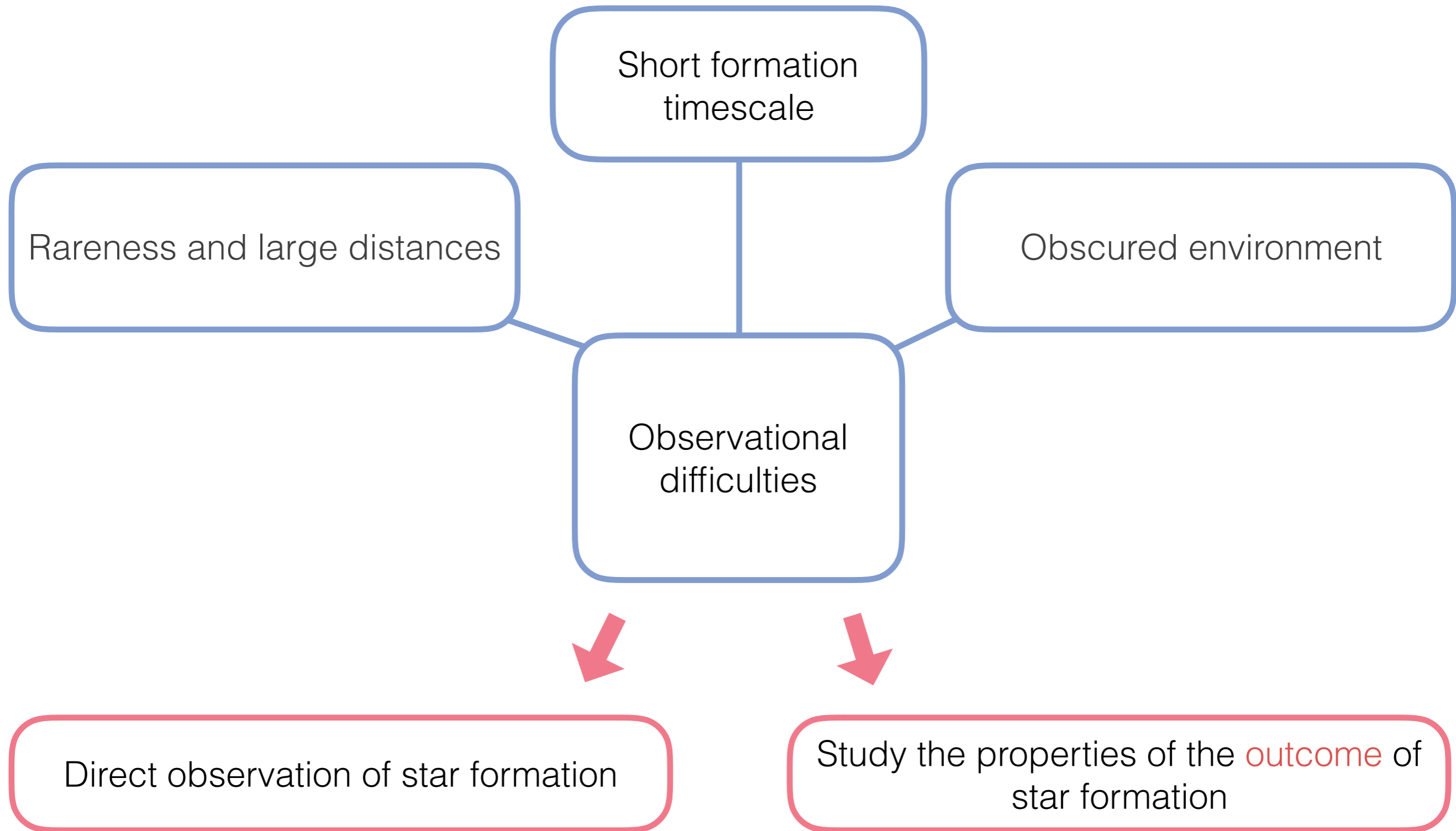
KU LEUVEN

What are massive stars?

Massive Star Formation



Massive Star Formation Problems



C

H

I

P

S

Carina

High contrast

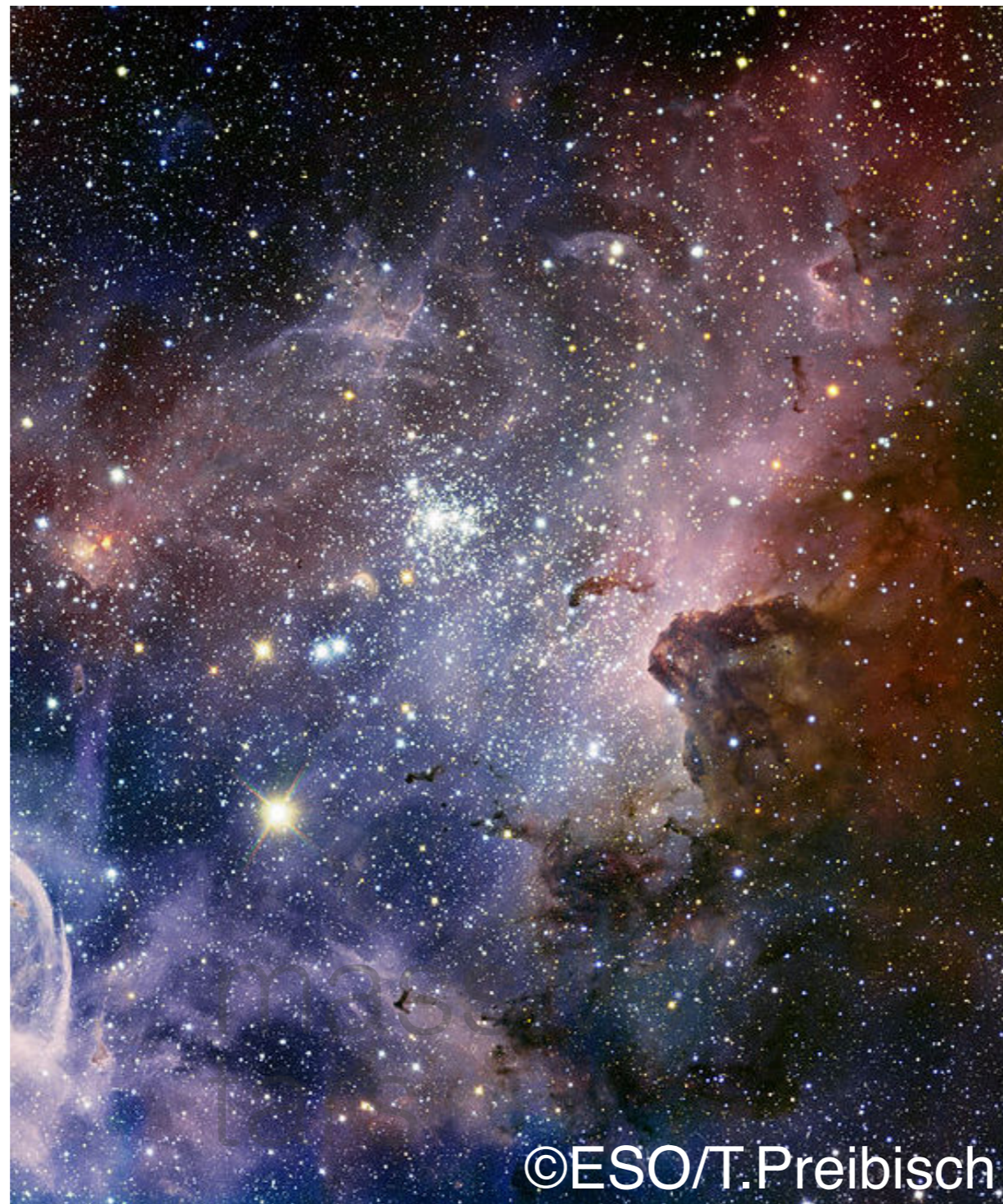
Imaging

Project of

Smassive
tars

Carina

close massive star
region



©ESO/T.Preibisch

Carina — close massive star region

High contrast
Imaging — VLT/SPHERE in IRDIFS mode

Project of
Smassive stars

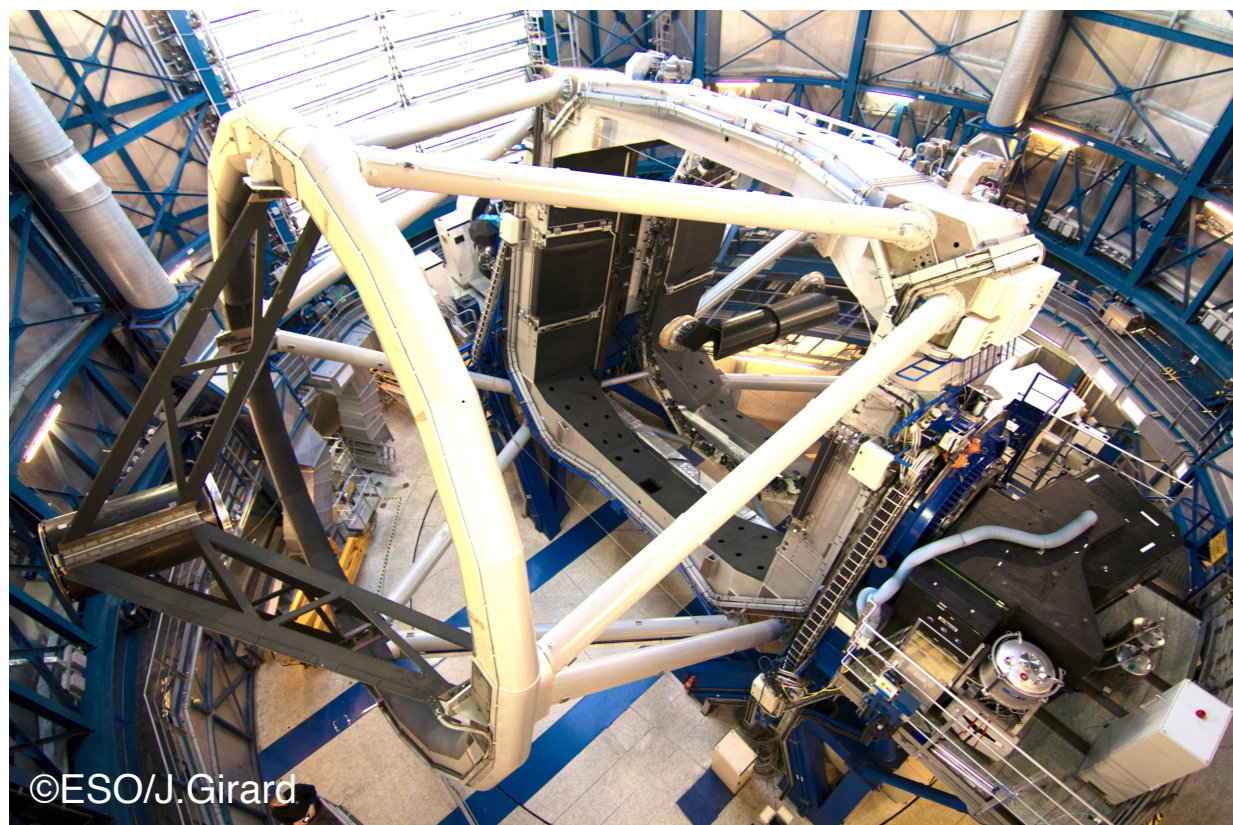
Carina

close massive star region

High contrast

VLT/SPHERE in IRDIFS mode

Imaging



©ESO/J. Girard

SPHERE	IFS	IRDIS
Spectral Range (μm)	0.95-1.75	0.95-2.32
FOV (arcsec^2)	1.73	11
Pixel Scale (marcsec)	7.4	12.25
Bands	Y-J-H	K (1&2)

Carina — close massive star region

High contrast
Imaging — VLT/SPHERE in IRDIFS mode

Project of — faint and low-mass companions

Smassive stars

Carina

close massive star

63 stars in P102

High contrast

28 stars already reduced

VLT/SPHERE in
IRDIFS mode

Imaging

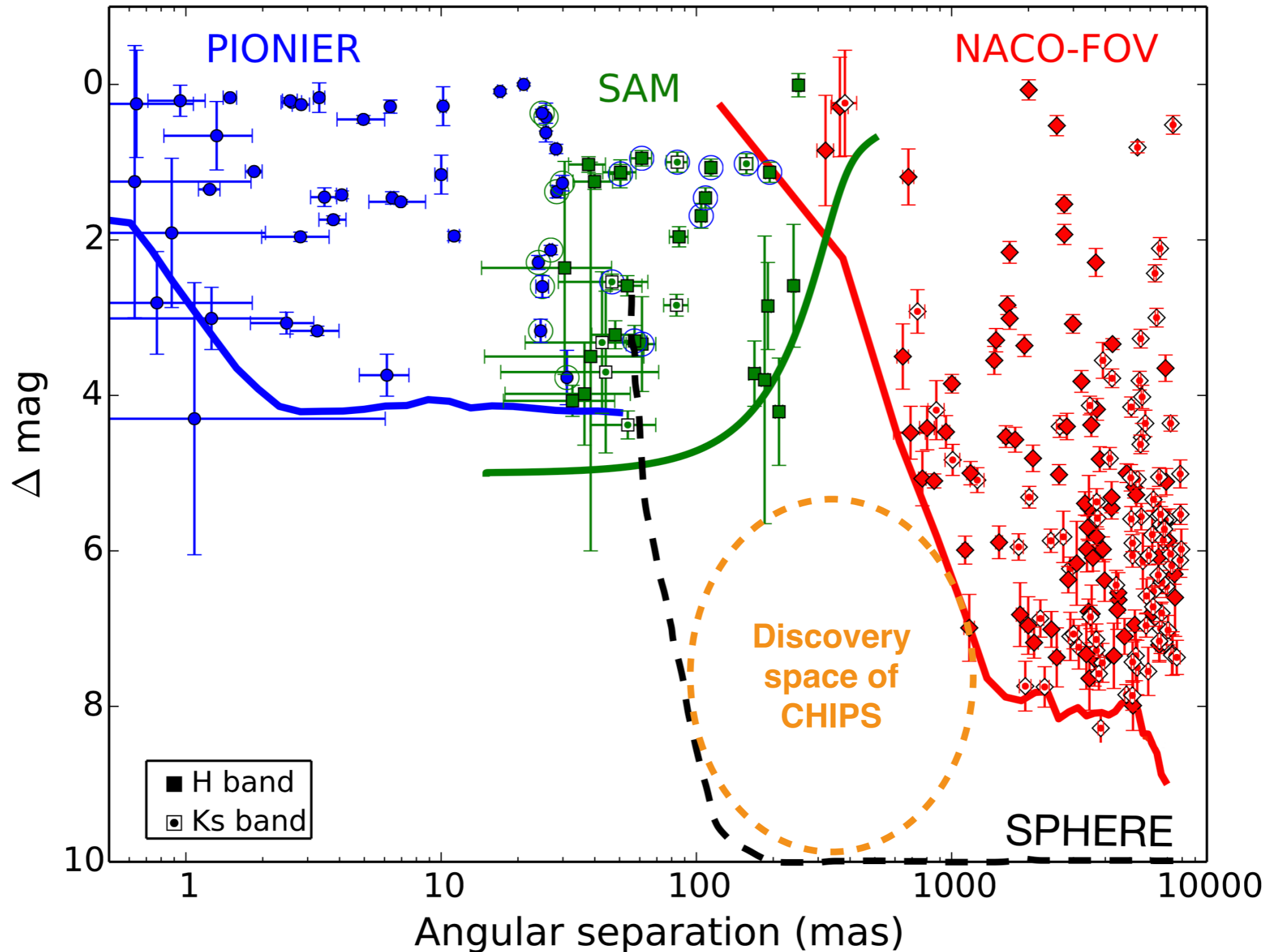
Project of

4D data cubes

S massive
stars

Multiplicity properties of
91 massive O and WR

SMaSH+ (Sana et al, 2014)

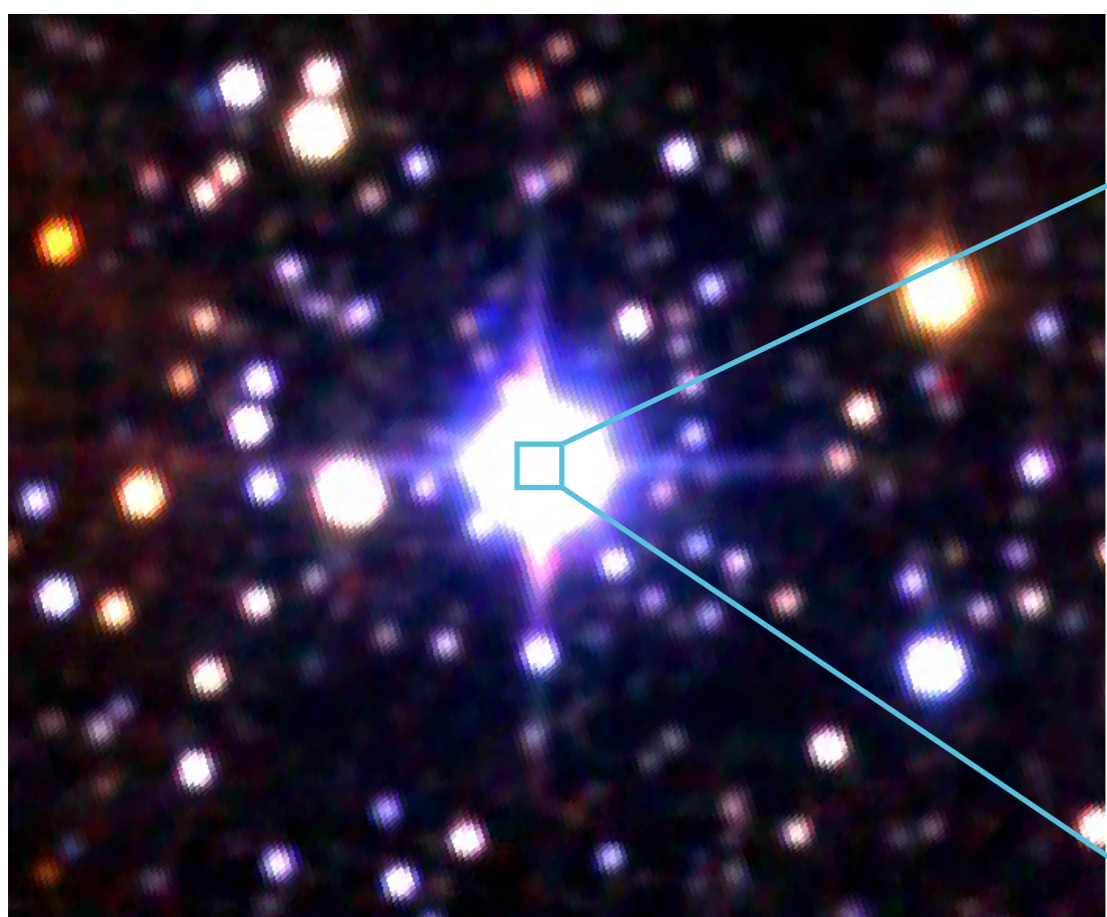


Some example data

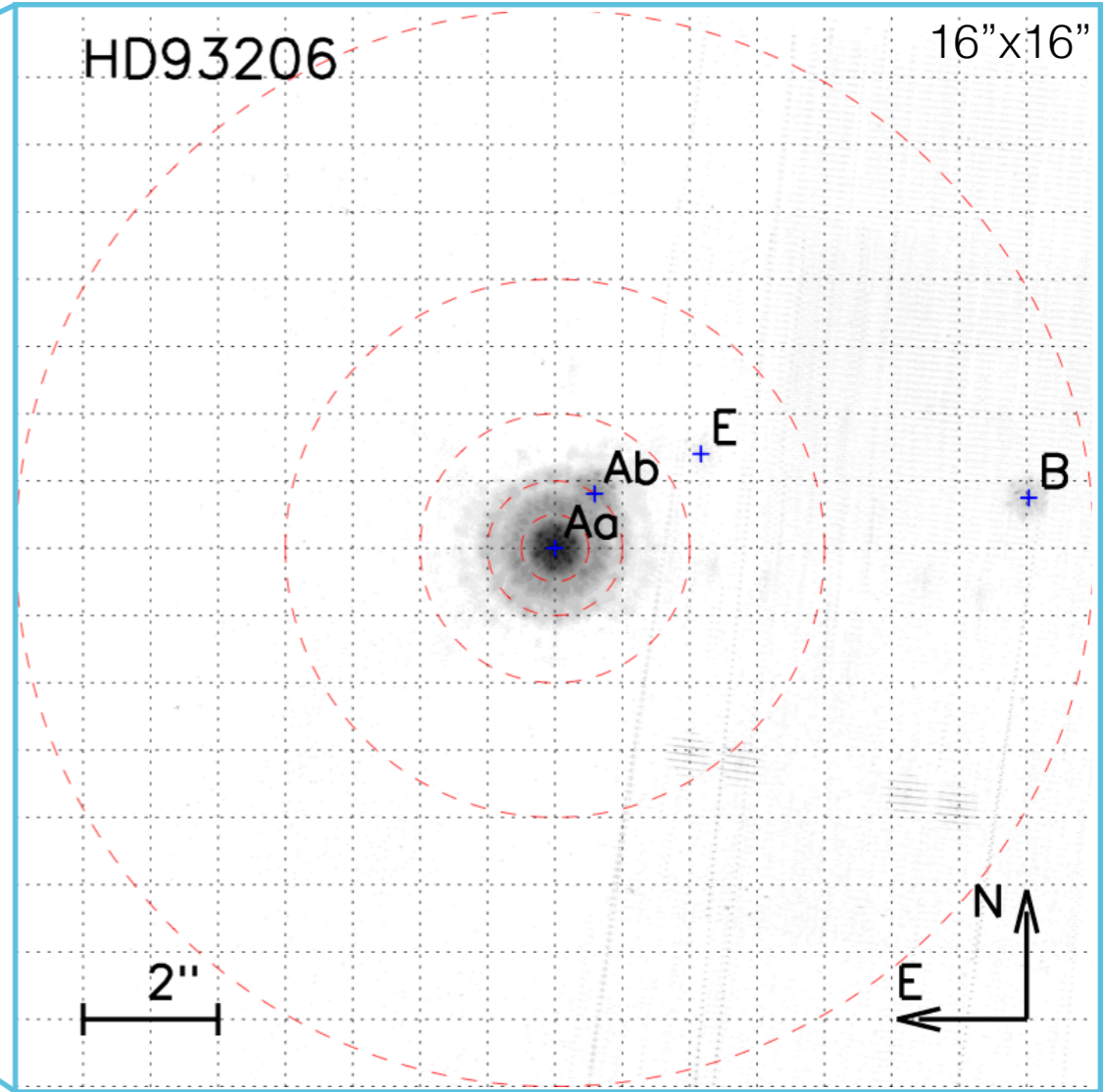
QZ Car (HD 93206)



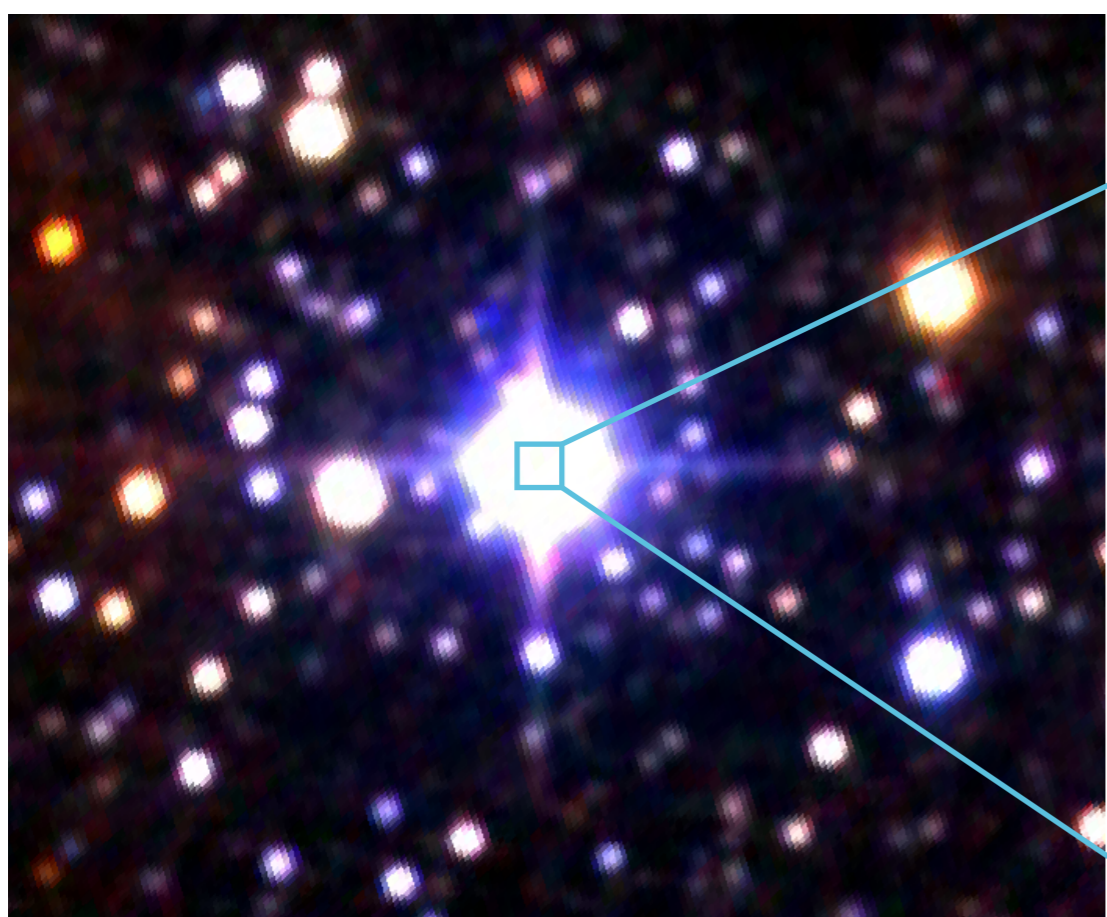
FOV = 5', 2MASS



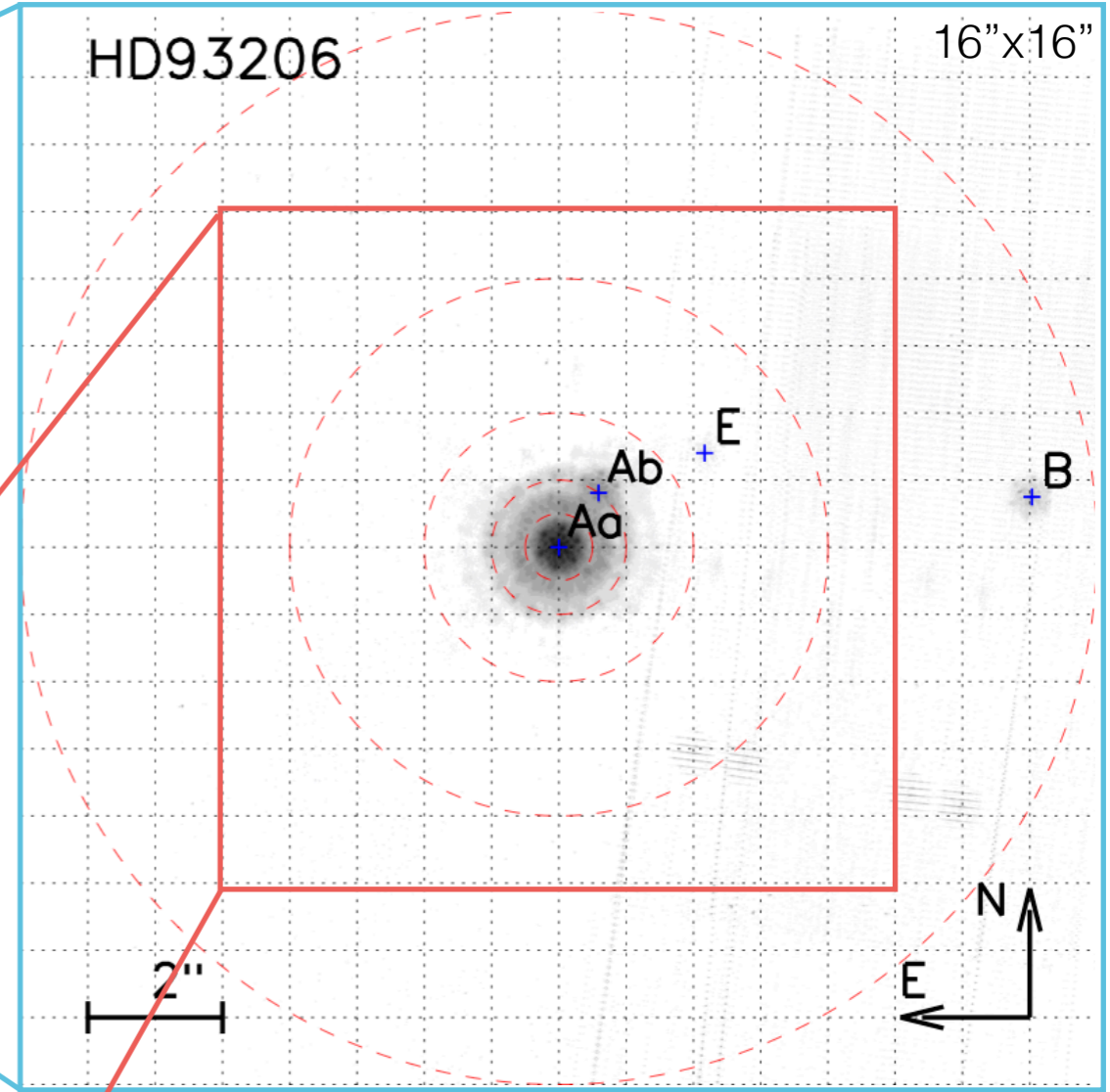
FOV = 5', 2MASS



SMA+ Image
Sana et al, 2014



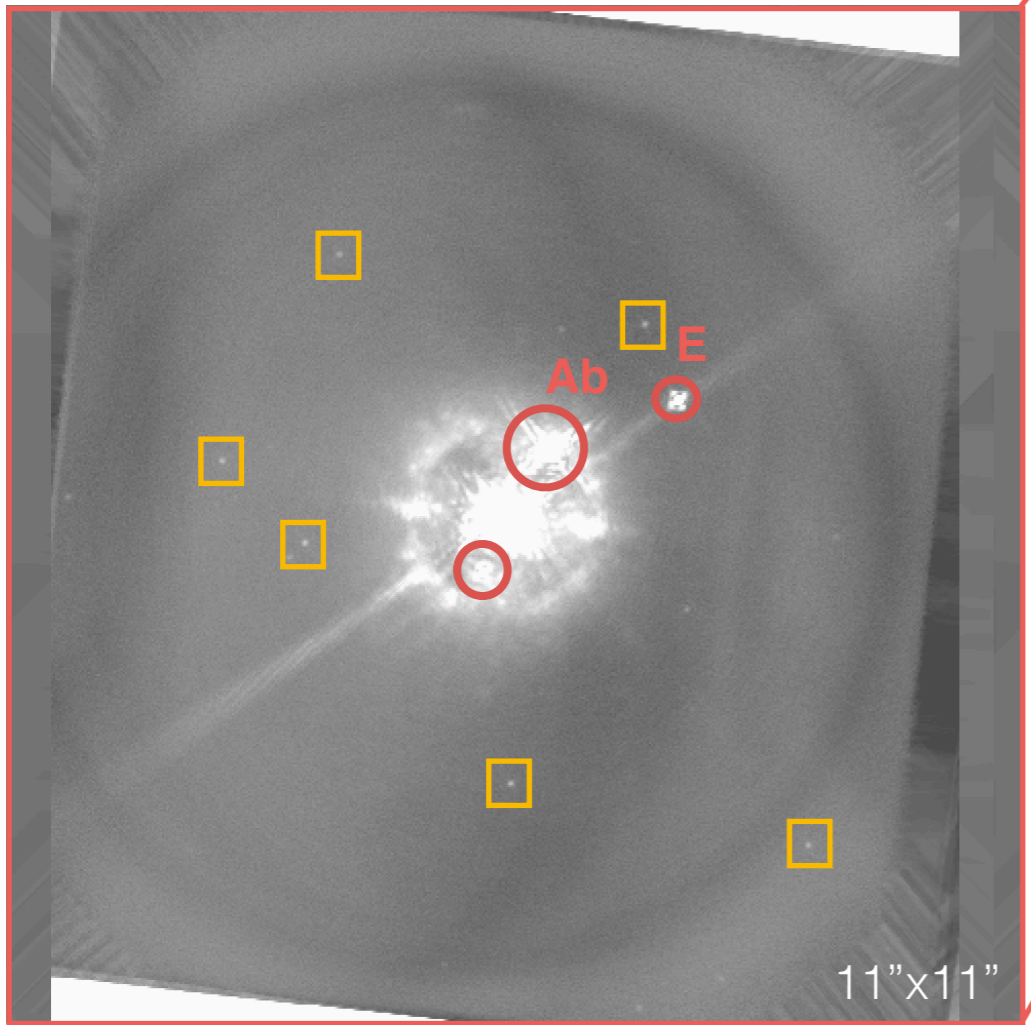
FOV = 5', 2MASS



HD93206

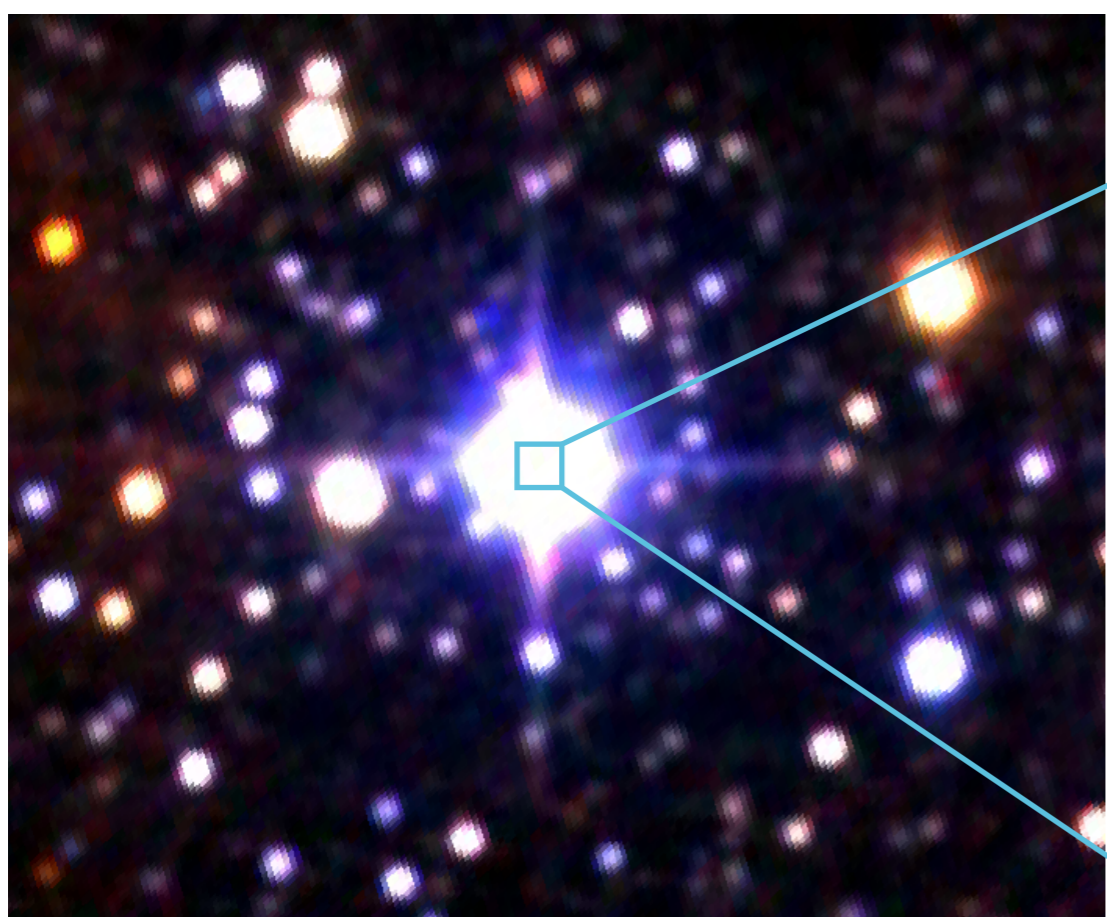
16"x16"

SMA+ Image
Sana et al, 2014

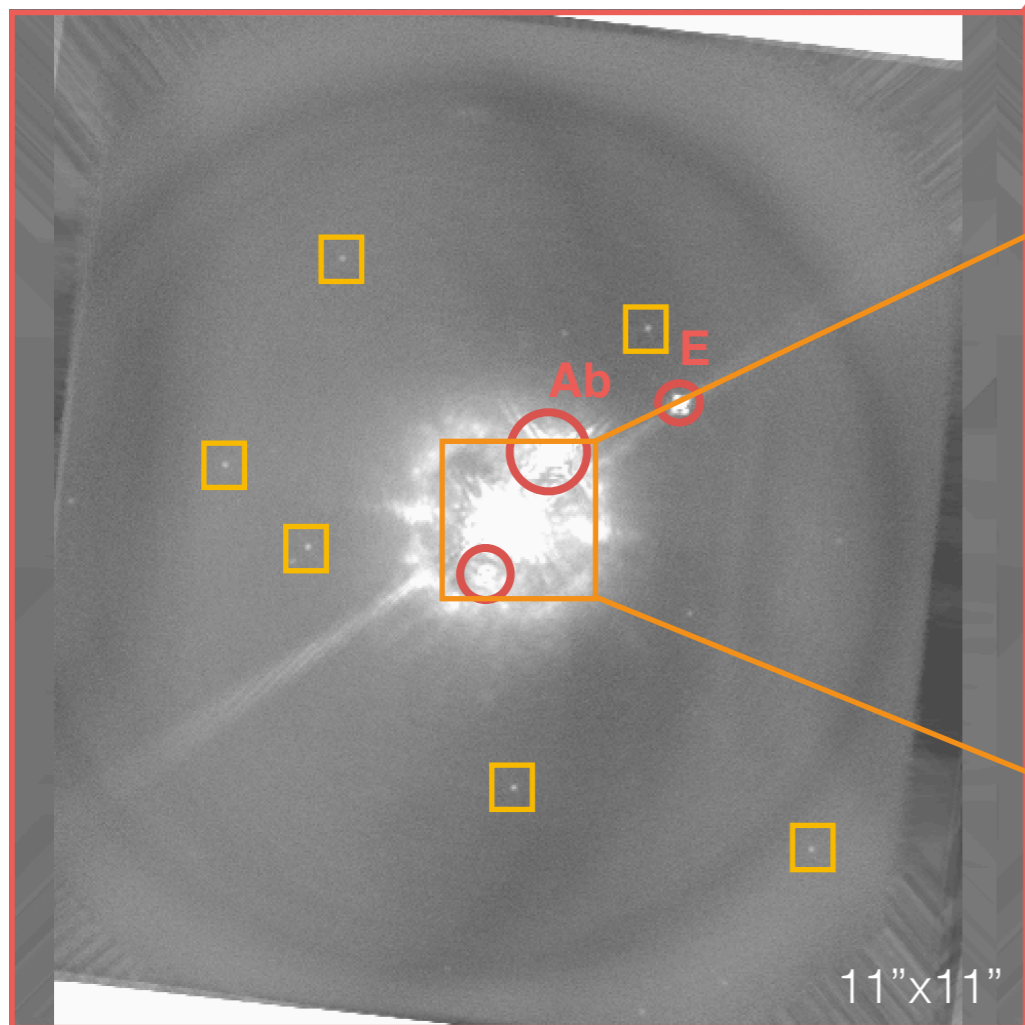
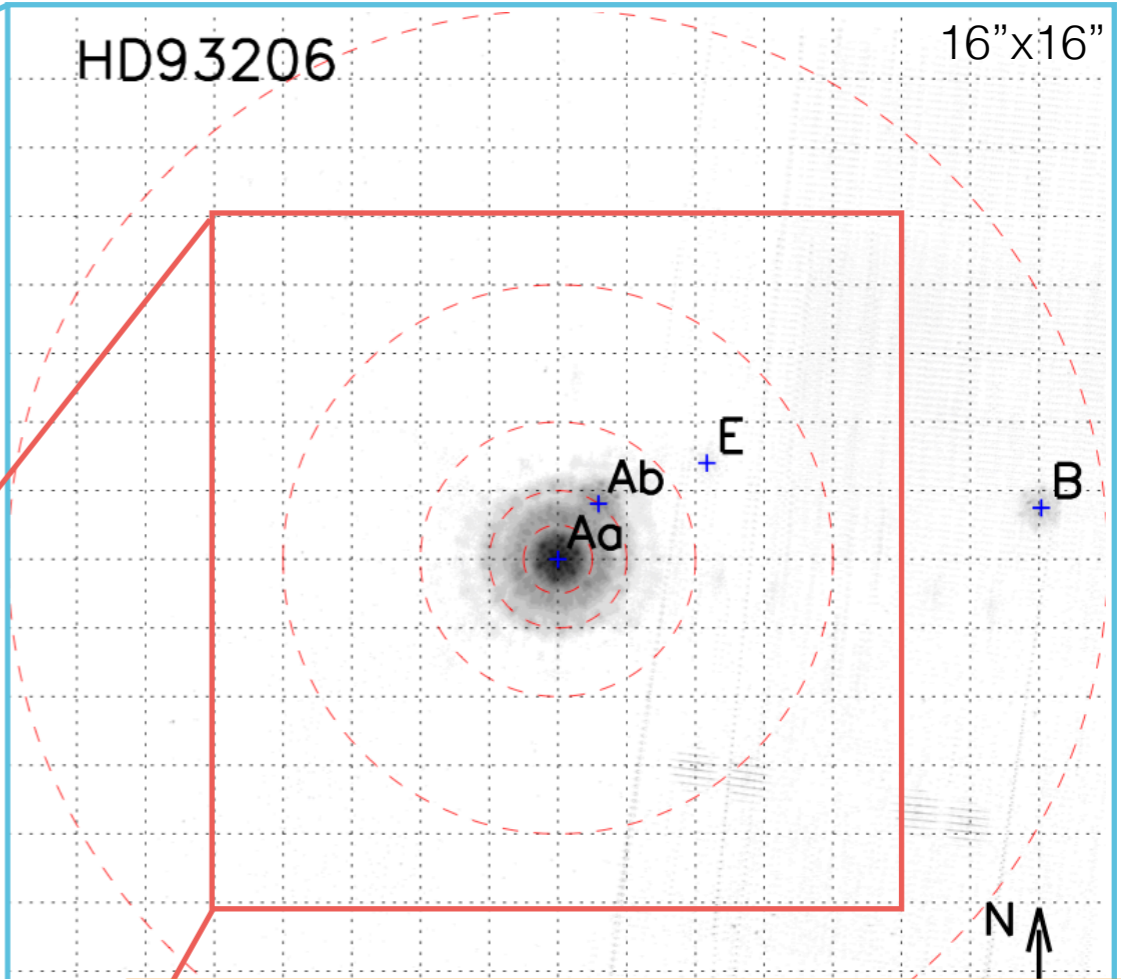


IRDIS

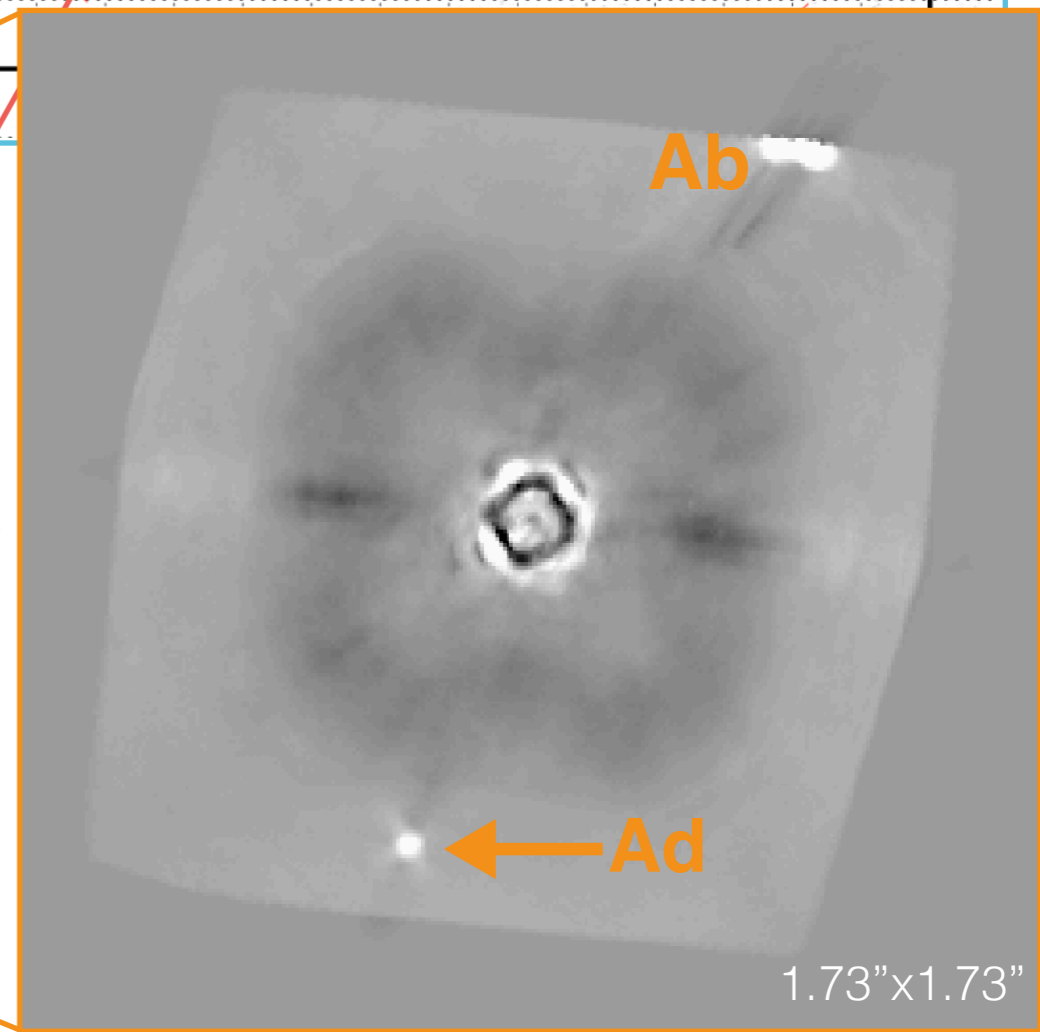
11"x11"



FOV = 5', 2MASS

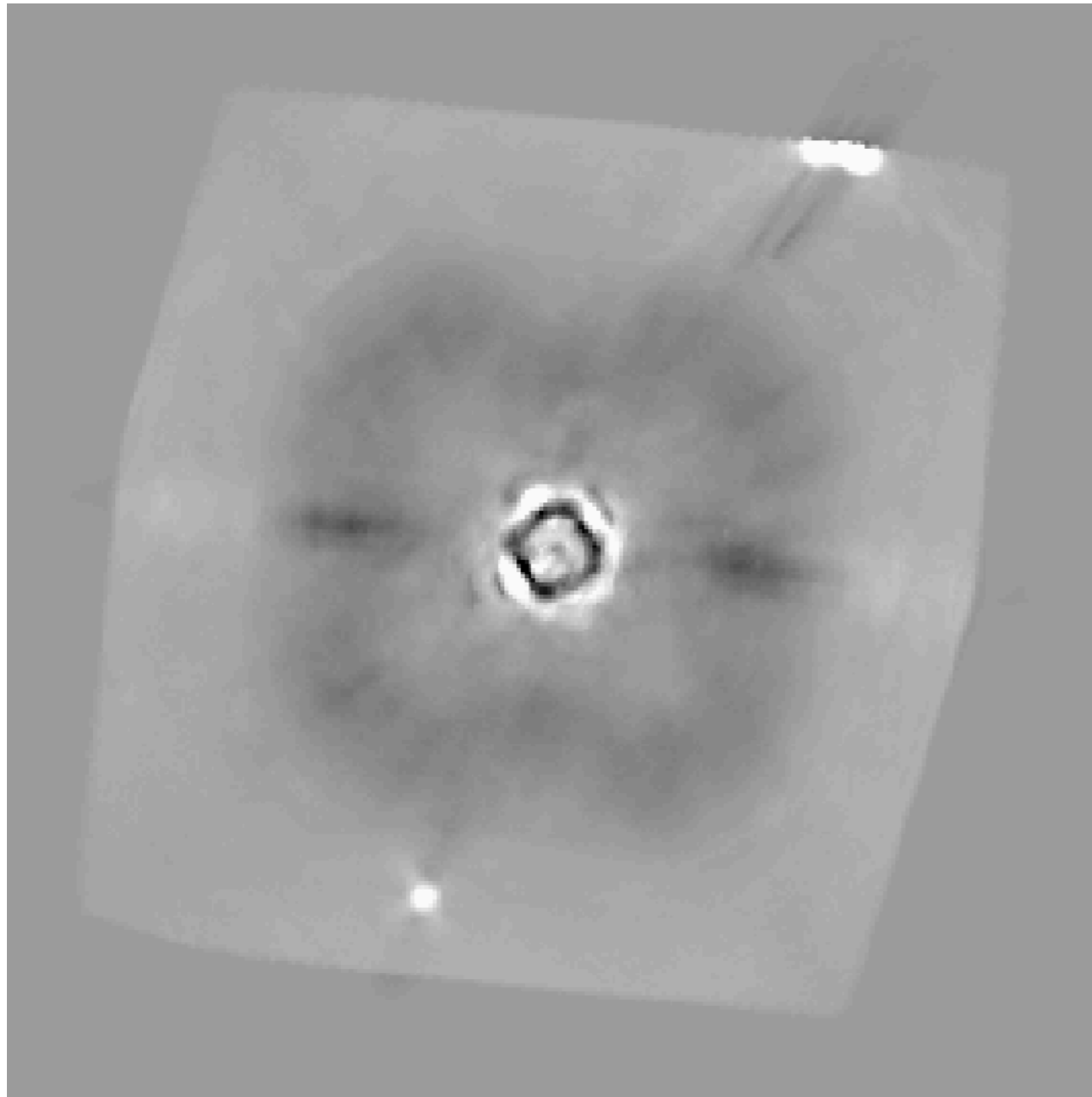


IRDIS

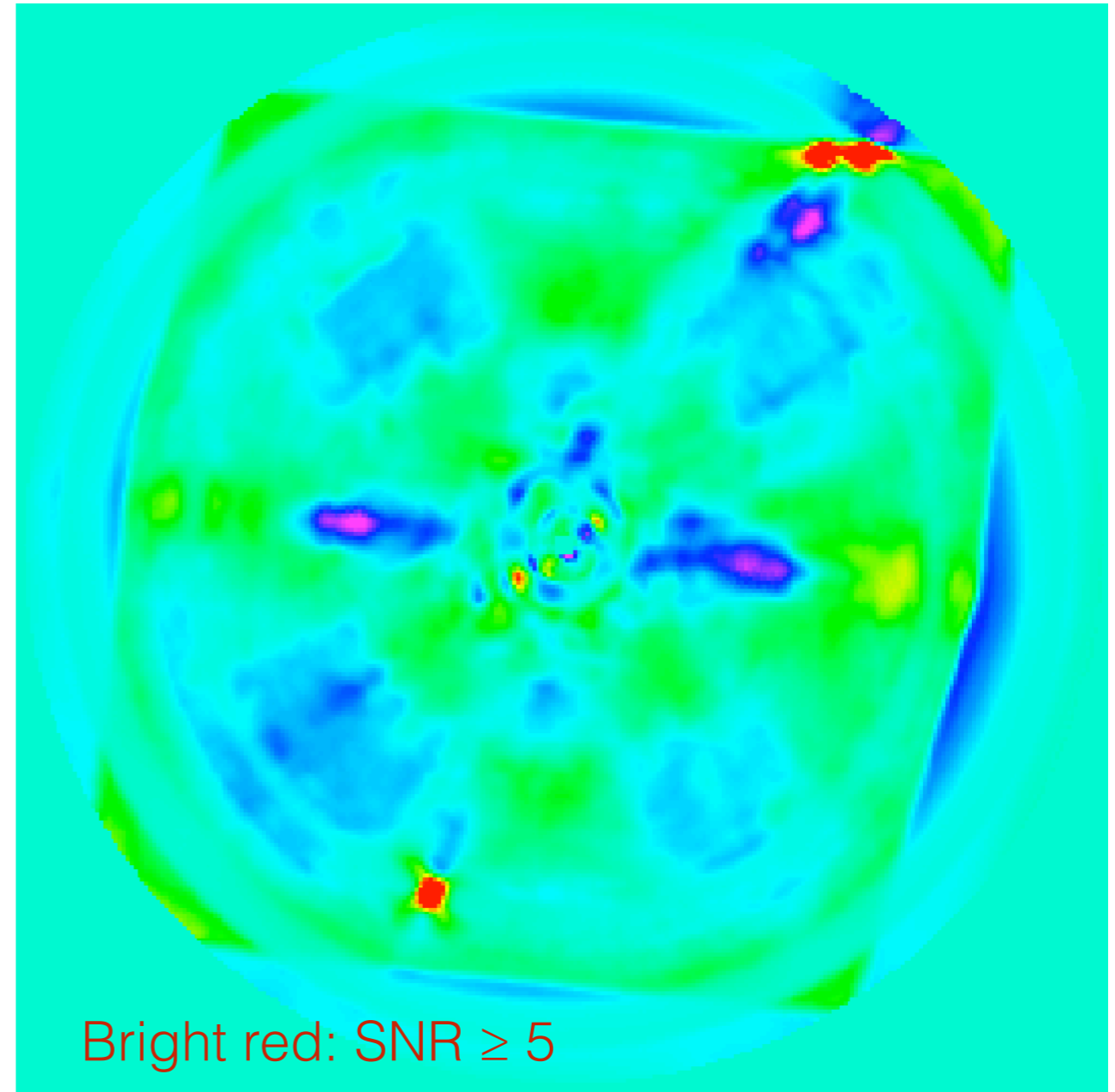


IFS

Detection method

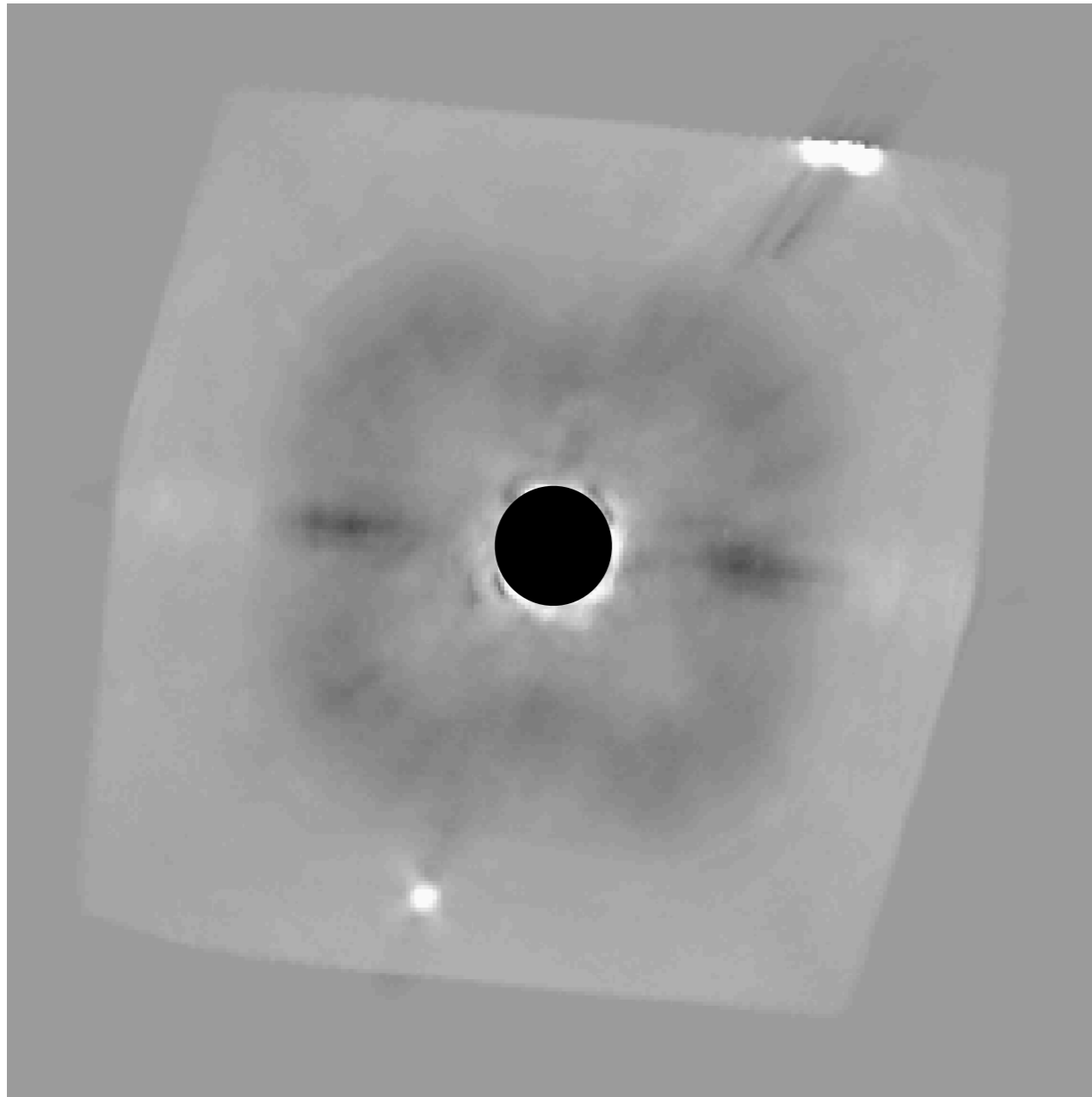


Derotated and wavelength collapsed image

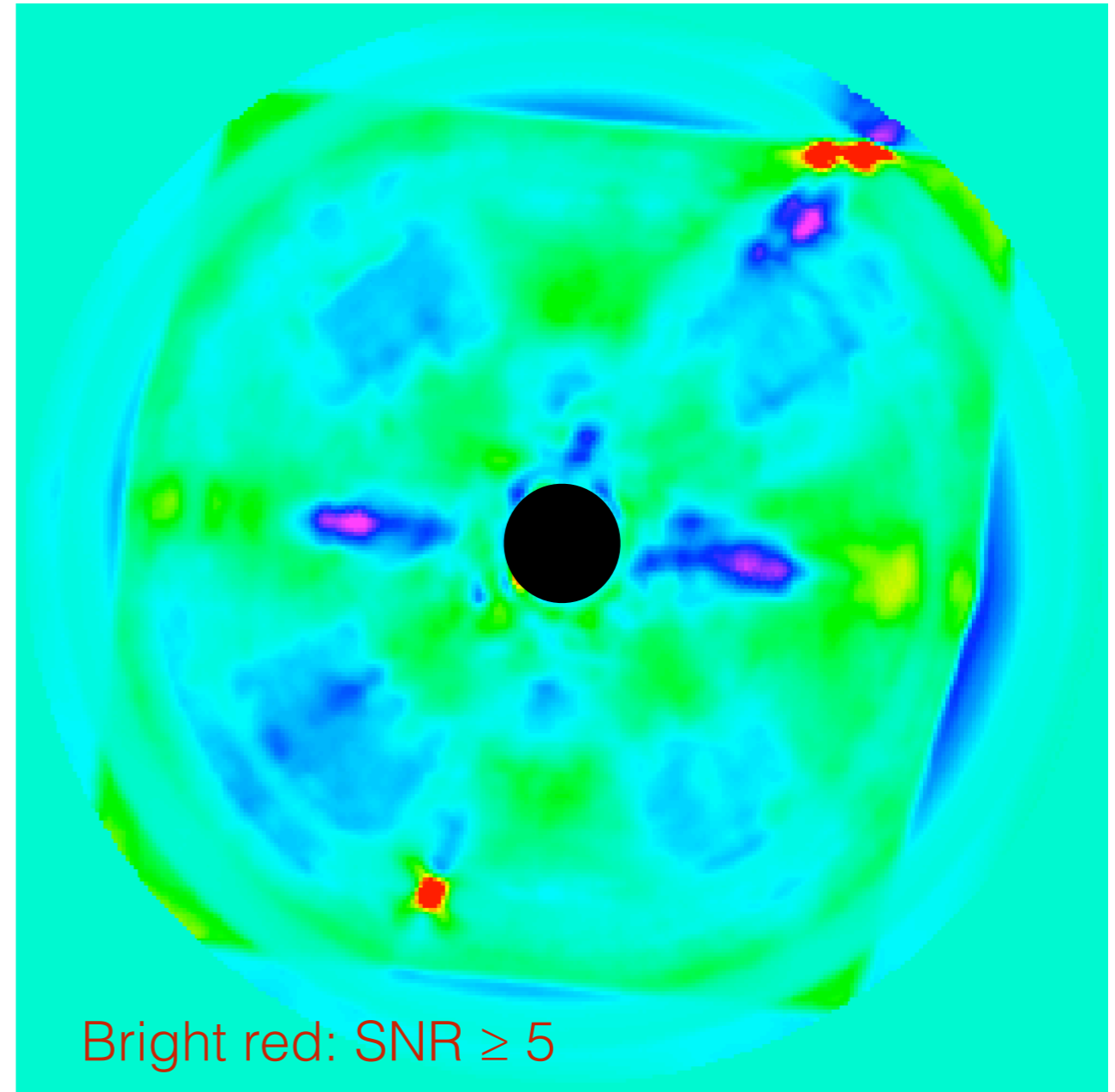


SNR map

Detection method

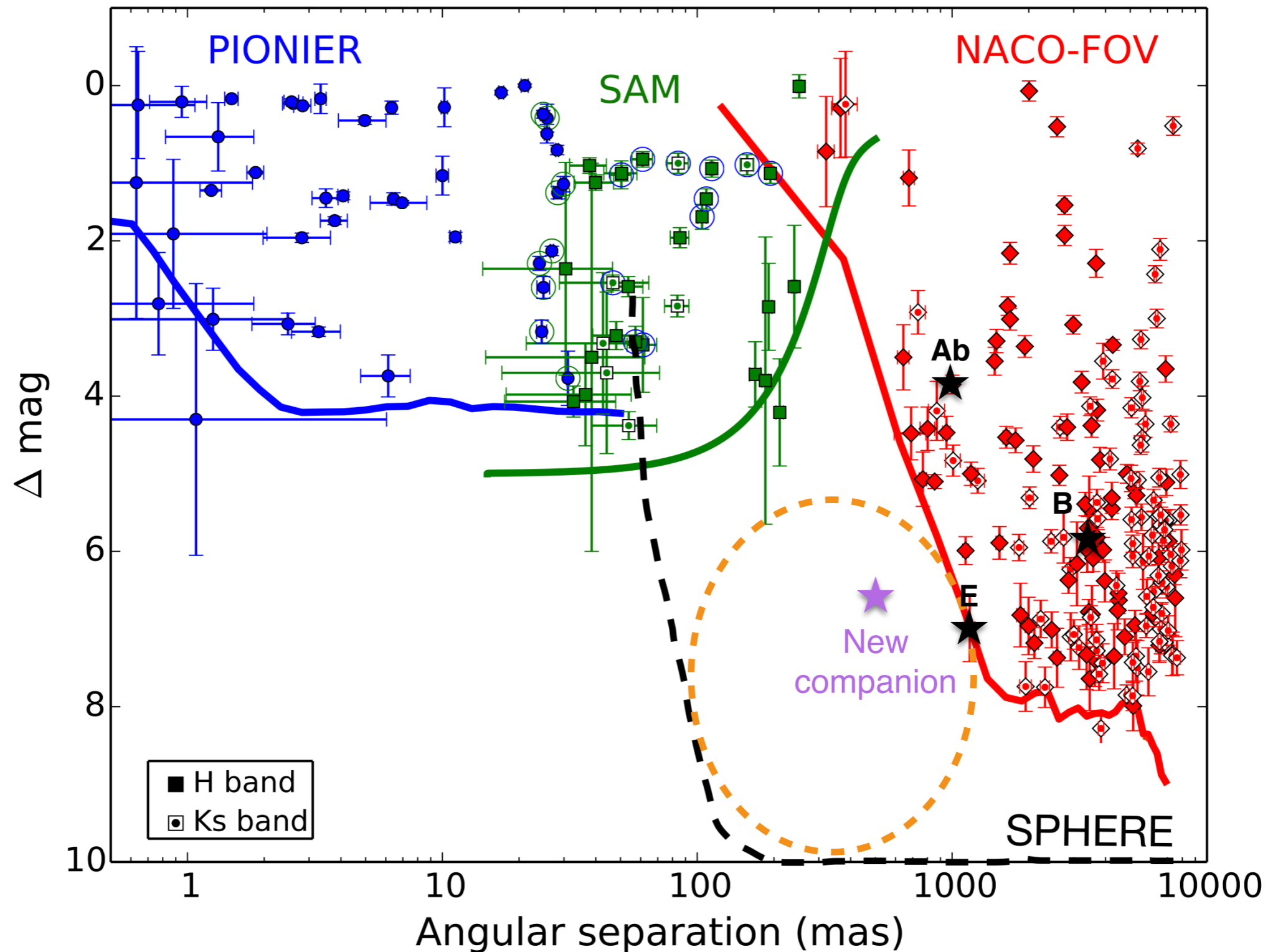


Derotated and wavelength collapsed image



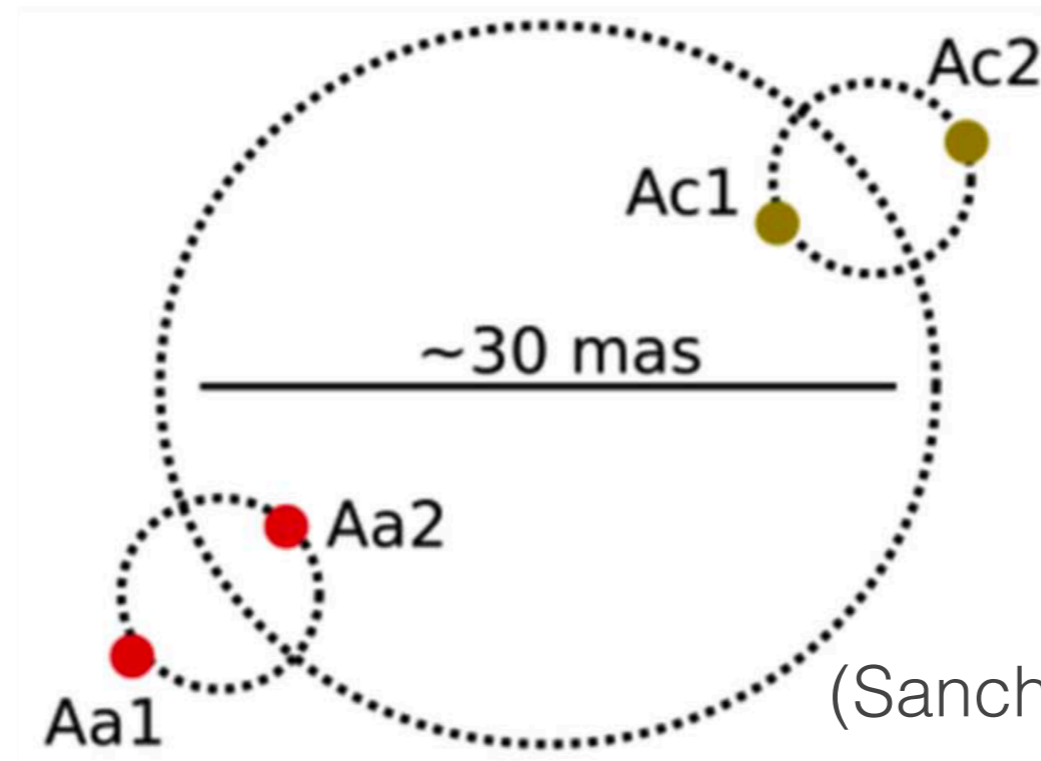
SNR map

SMaSH+ (Sana et al, 2014)



Spectrum Extraction

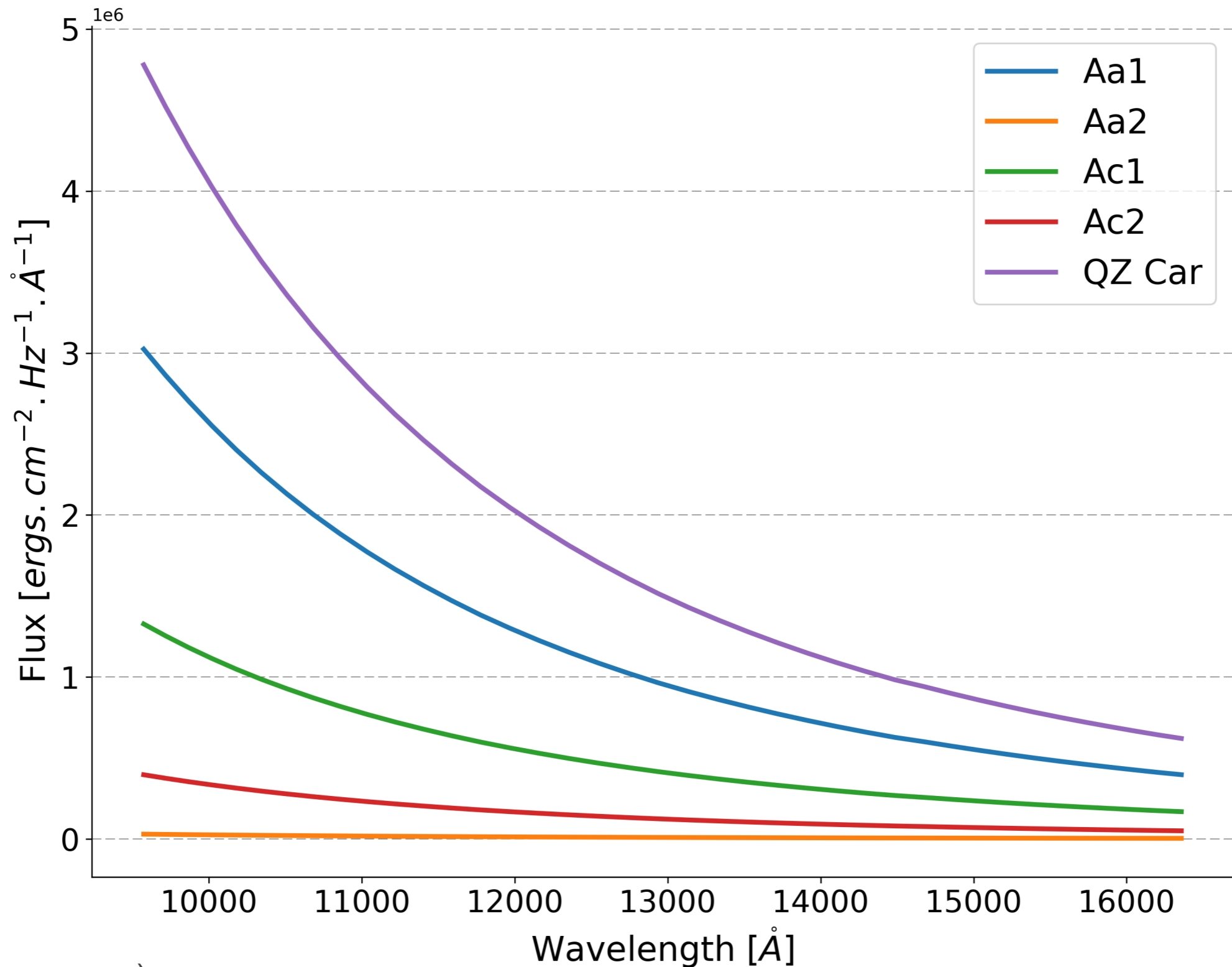
QZ Car (HD 93206)



Component	Spectral Type	T_{eff} (K)	R_* (R_\odot)	M_* (M_\odot)	L_* ($\log[L_*/L_\odot]$)	\dot{M} ($M_\odot \cdot yr^{-1}$)	$\log(g)$	v_∞ ($km \cdot s^{-1}$)
Aa1	O9.7 I	32000	22.5	40	5.7	8.21×10^{-6}	3.19	1794.3
Aa2	b2 v	20000	6.0	10	3.7	2.39×10^{-14}	4.3	1186.4
Ac1	O8 III	32573	26.9	14.1	5.3	3.32×10^{-6}	3.57	2191.2
Ac2	o9 v	32463	8.9	28	4.9	3.16×10^{-9}	3.92	2427.1

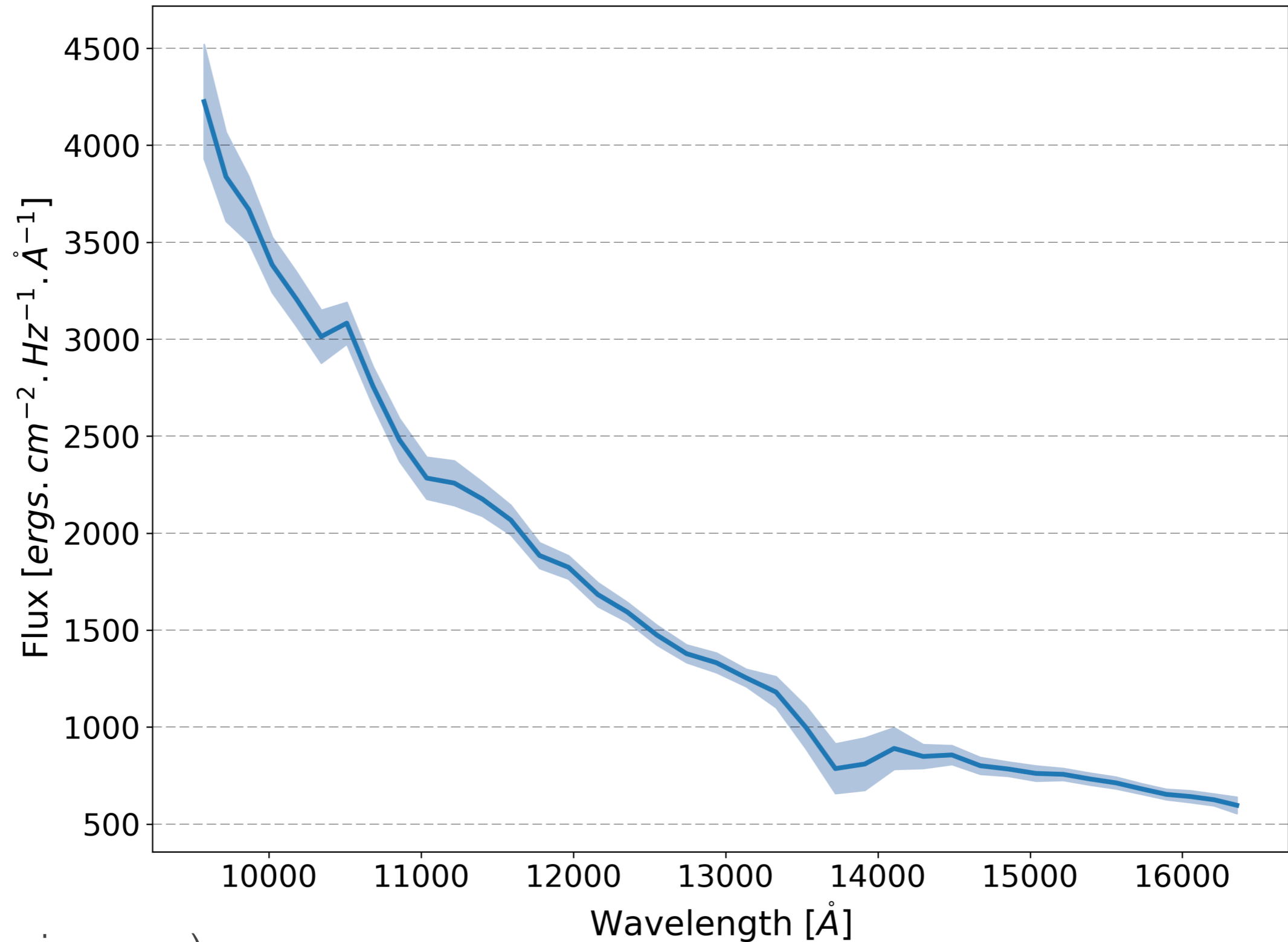
(Rainot+ in prep.)

QZ Car spectrum with FASTWIND



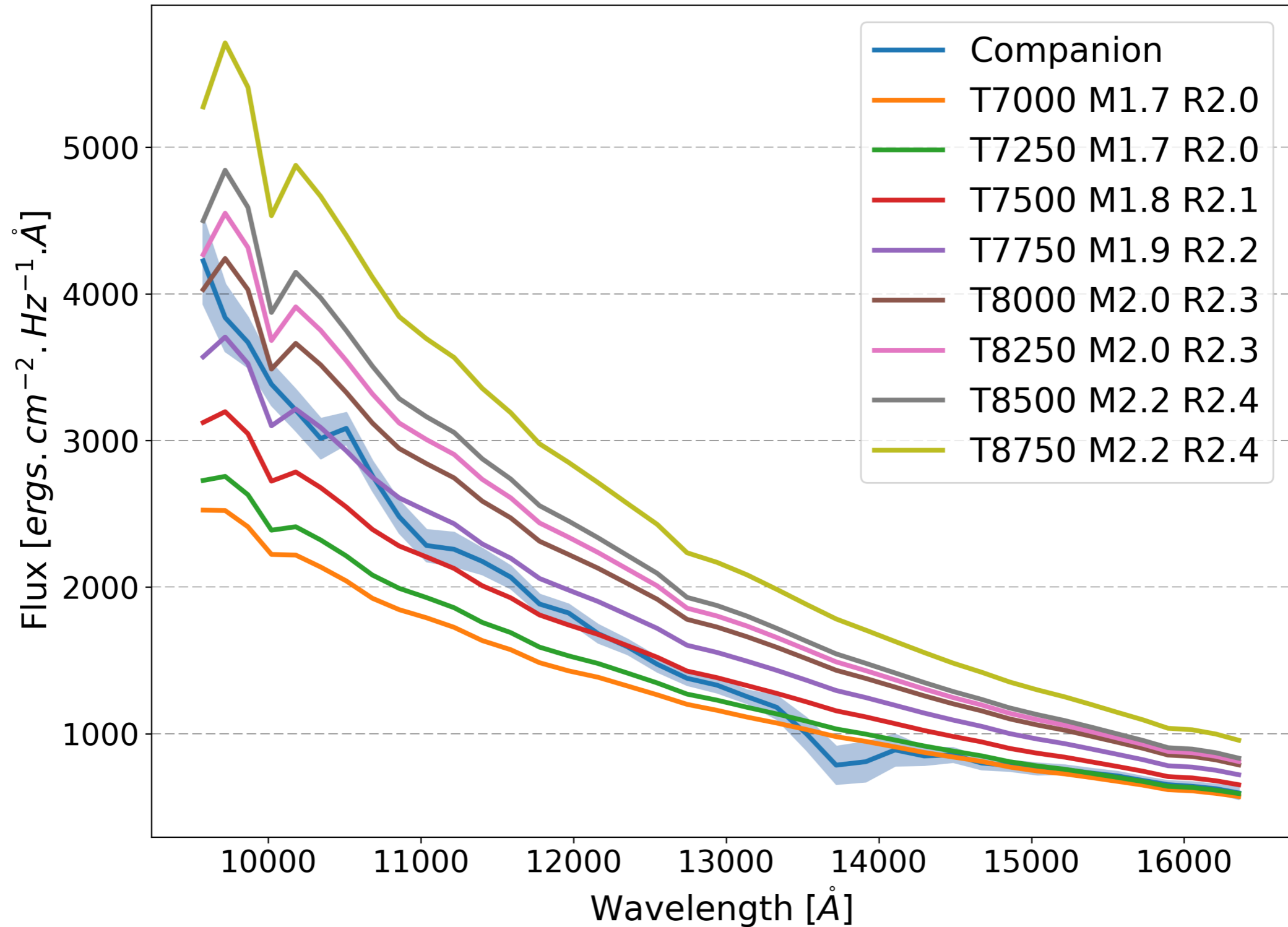
(Rainot+ in prep.)

Calibrated spectrum



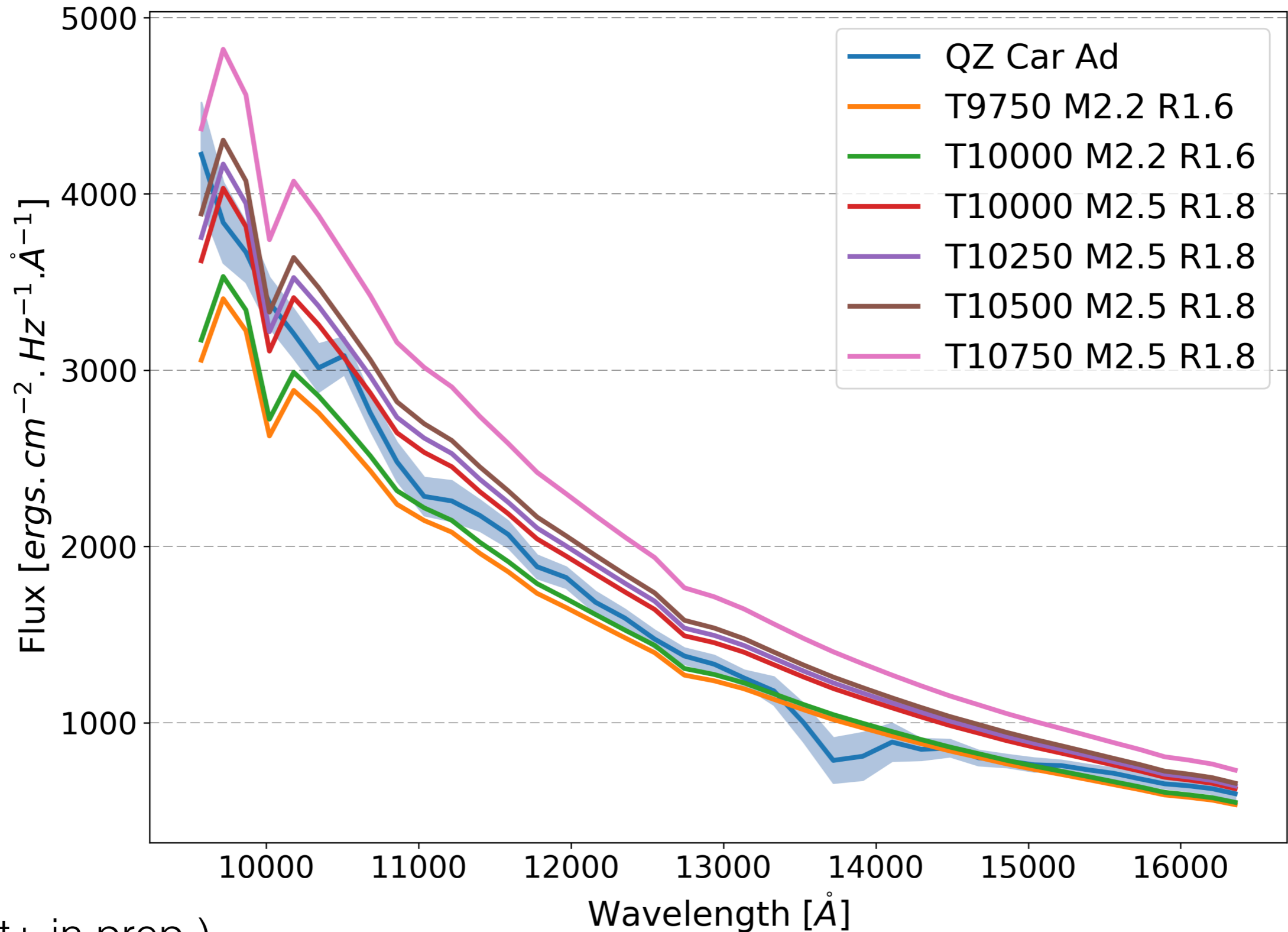
(Rainot+ in prep.)

ZAMS



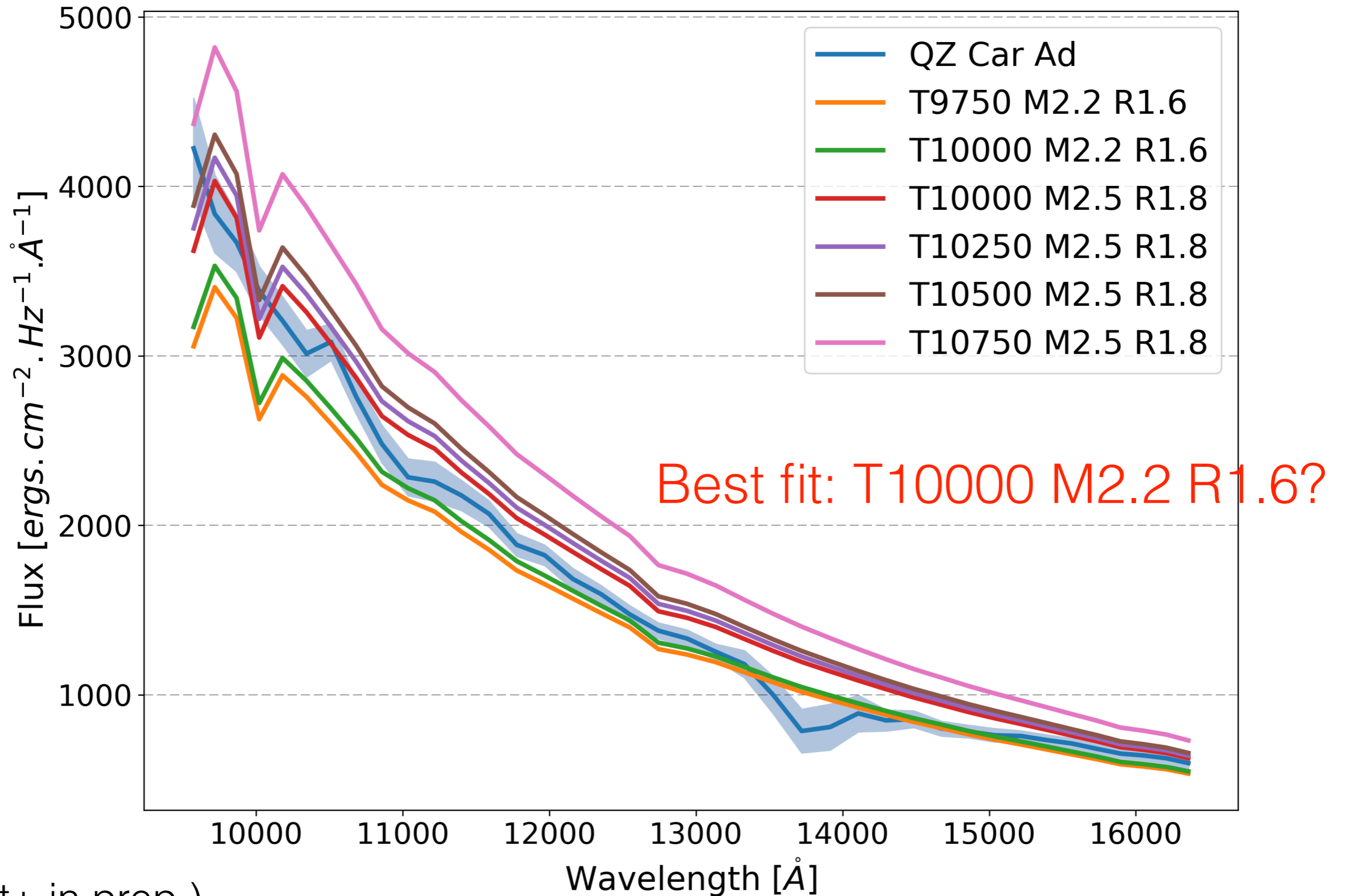
(Rainot+ in prep.)

Early MS



(Rainot+ in prep.)

Early MS



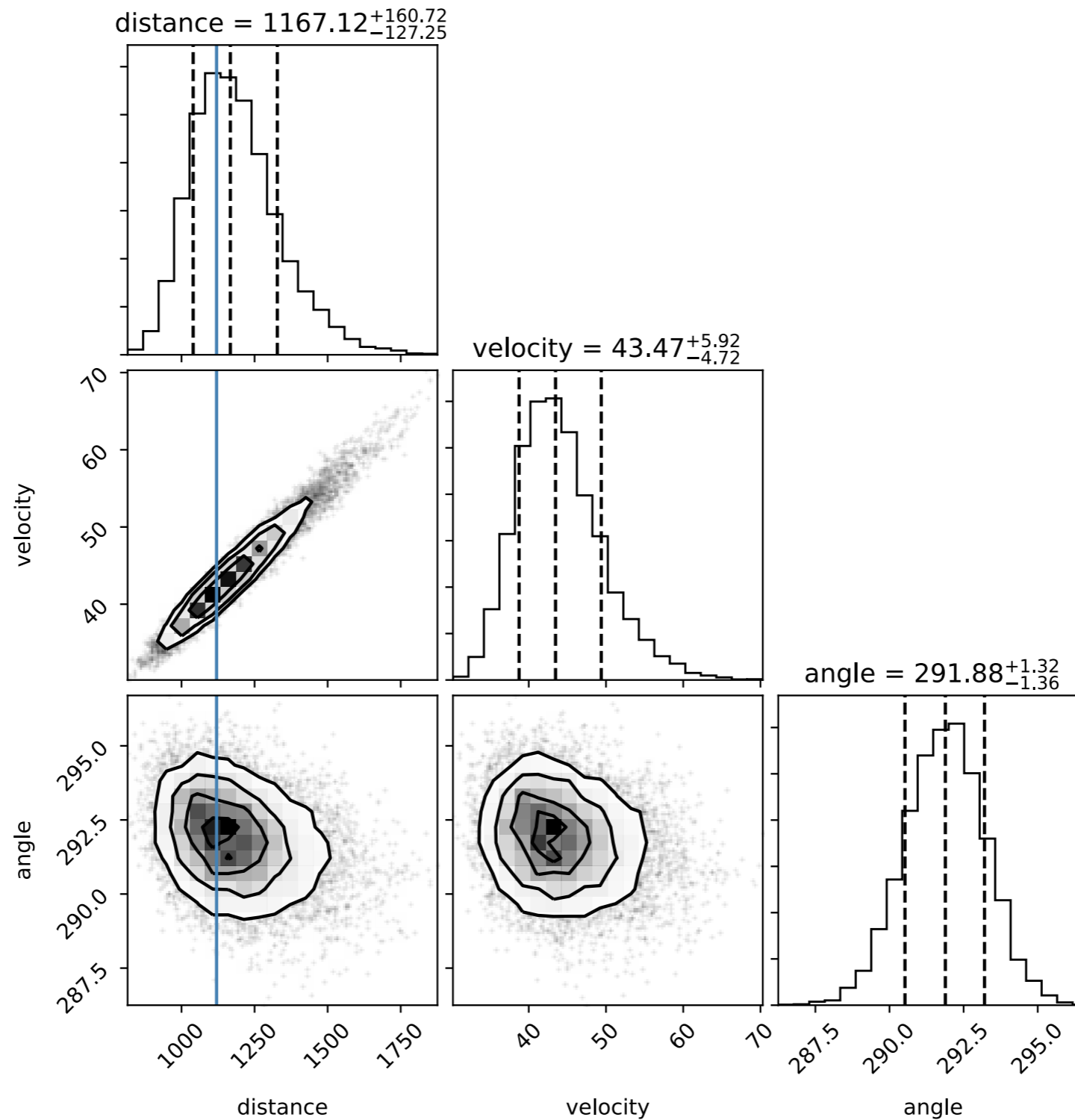
(Rainot+ in prep.)

Multiplicity Results - IFS

Images	Visible companions ($\text{SNR} > 5\sigma$)	Candidates ($5\sigma > \text{SNR} > 3\sigma$)
28	6	12

- ▶ Detection ratio ≈ 0.42 companions/star
- ▶ Expected ≈ 38 companions / 91 images

GAI A distance



(Rainot+ in prep.)

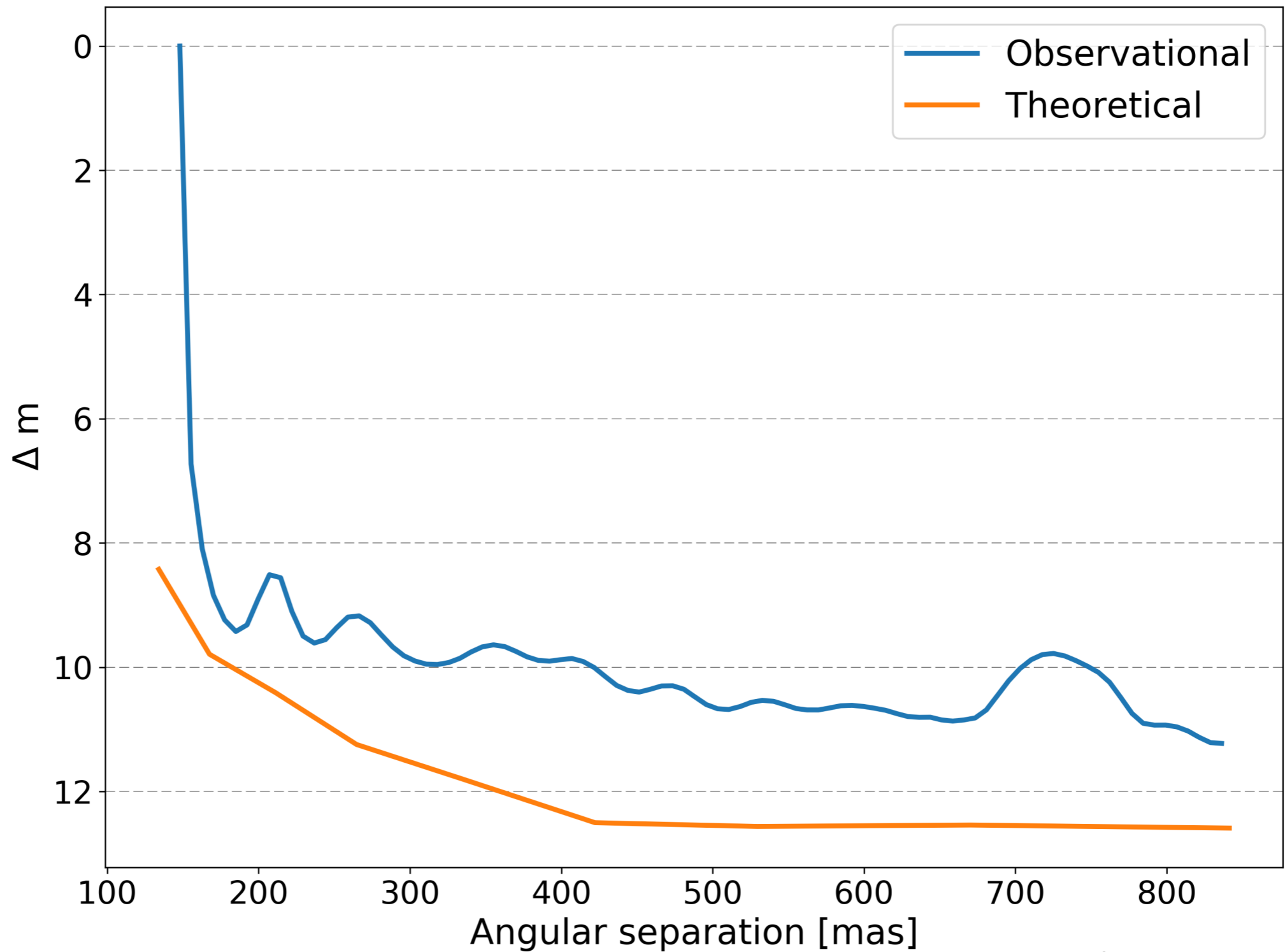
Conclusion

- ▶ **Discovery** of a new companion: QZ Car Ad at 1.2kpc
- ▶ Might be an **A0 star** with: $T = 10000\text{K}$, $M = 2.2M$, $R = 1.6R$
- ▶ Masses limits of IRDIS **detected sources**
- ▶ **63 stars** remain to be observed with SPHERE (P102)
- ▶ First paper to be submitted **soon!**

Thank you!

Extra Slides

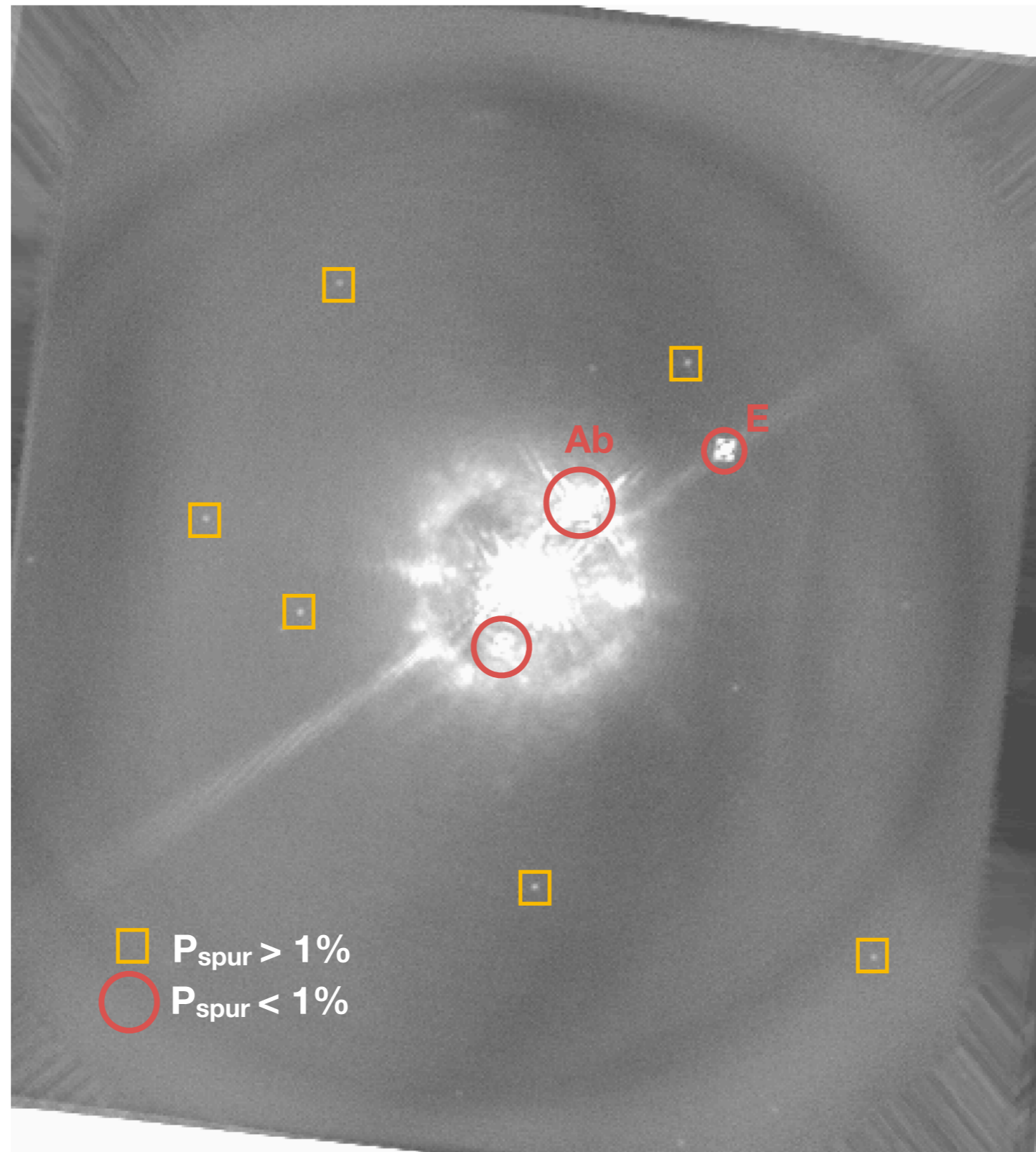
IFS Contrast Curves



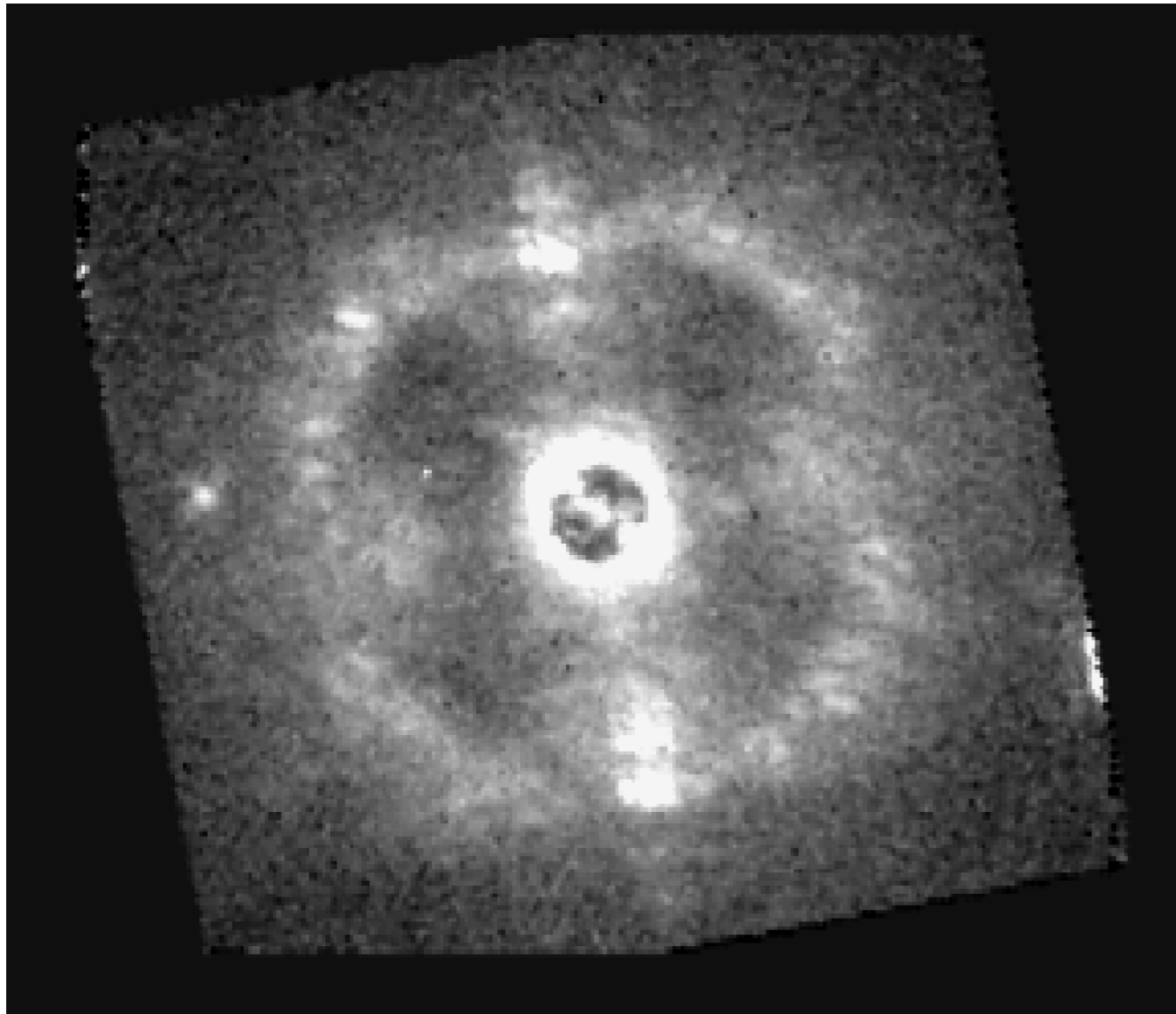
(Rainot+ in prep.)

IRDIS

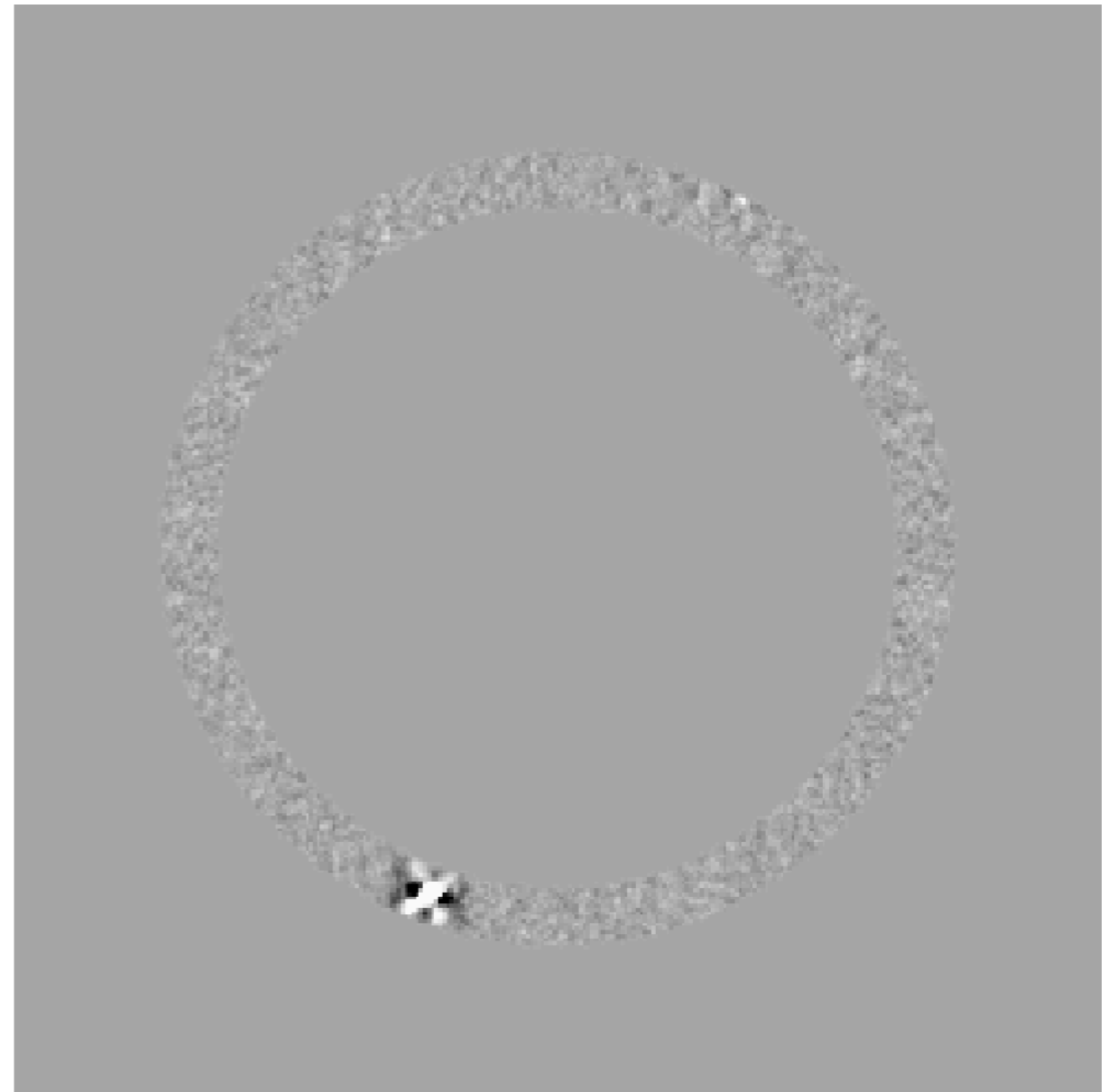
(Rainot+ in prep.)



Analysis Techniques



Original image



Post-processed cube

PSF Fitting

